



Remote sensing of fires with Meteosat



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IMAGES

REAL-TIME IMAGES

IMAGE LIBRARY

IMAGE GLOSSARY

FILTER BY

Feature

Fire

Drought

Dust

Equinox

Extratropical transition

Fire

Flood

Föhn

Fog

Front

Frost

Gap Wind

Gravity Wave

Gust Front

Haboob

Hail

Haze

Heatwave

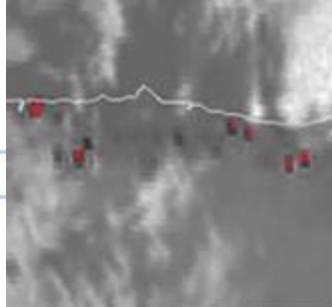
Hot Spot

HRV Winds

Humidity

▶ APPLY

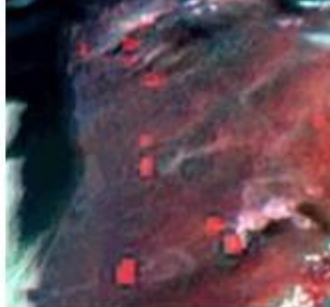
▶ RESET



FIRES DEVASTATE PARTS OF ASTURIAS AND CANTABRIA

19 December 2015

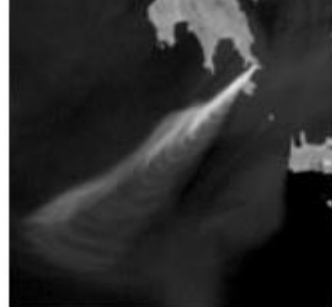
Northern regions of Asturias and Cantabria in Spain were ravaged by hundreds of fires in December.



FIRES IN THE IBERIAN PENINSULA

06 August 2015

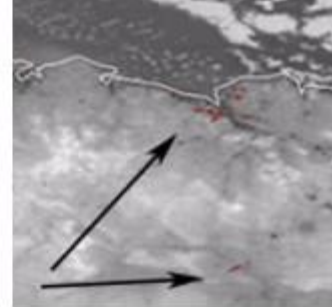
Due to the hot summer and consequent dry soil, there were a number of fires in the Iberian peninsula, on 6-11 August.



WILDFIRES IN GREECE

17 July 2015

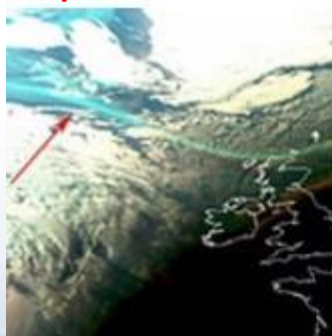
In the night of 16/17 July wildfires broke out in the extreme south-eastern corner of the Peloponnese peninsula.



MIDSUMMER FIRES

23 June 2015

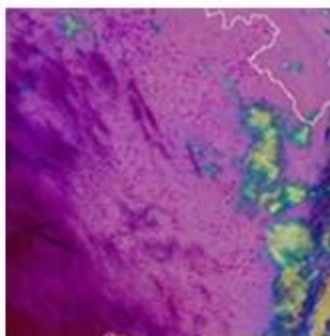
Large Midsummer fires in Spain, on the evening of 23 June, were detected by Metop-A's 3.7µm channel.



SMOKE FROM CANADIAN FIRES SEEN OVER EUROPE

13 June 2015

On 13 June smoke from wildfires in Canada travelled over parts of Northern Europe.



TRADITIONAL BELTANE FIRES SEEN BY METOP

30 April 2015

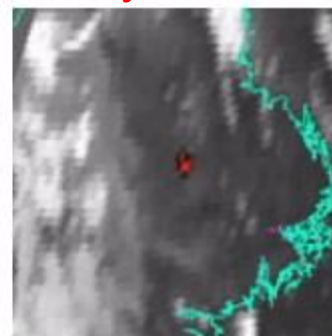
Hotspots from balefires in parts of Europe could be seen on the Metop image from 30 April.



FIRES OVER THE SOUTH-WESTERN PARTS OF SOUTH AFRICA

12 February 2015

Many runaway/out of control fires over the south-western parts of



LARGE FOREST FIRE IN SWEDEN

04 August 2014

Meteosat and Suomi-NPP imagery of a large forest fire in Sweden showing hot spots with temperatures up to 356 K.

<http://www.eumetsat.int/website/home/Images/ImageLibrary/index.html>

Forest Fires

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The Fire Triangle

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The Continuity Equation

Fuel

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EUMeTrain

As you have seen in the previous chapter, synoptic patterns are not enough to indicate the locations where forest fires are likely to occur and spread. In fact, aside from the weather, a fire's behaviour also strongly depends on fuel and topography, creating the three sides of what is usually known as the *fire triangle* (Figure 3.1).

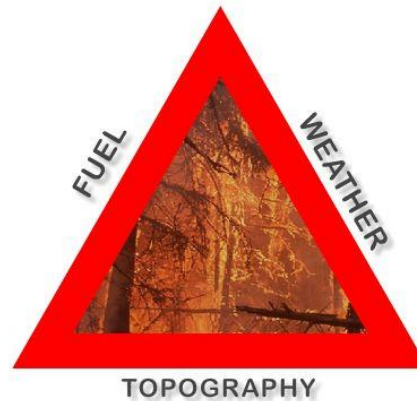


Fig 3.1 – Fire triangle. Fire hazards are influenced by fuel, weather and topography.

In this third chapter the three sides of this triangle will be further explained. After answering the following multiple choice question the chapter will continue with "weather".

Fire ignition depends most on which of the following meteorological factors:

- ☐ a) Temperature



Characteristics of SEVIRI channels

- Characteristics of the 3.9 μ m channel
- Solar channels for monitoring vegetation

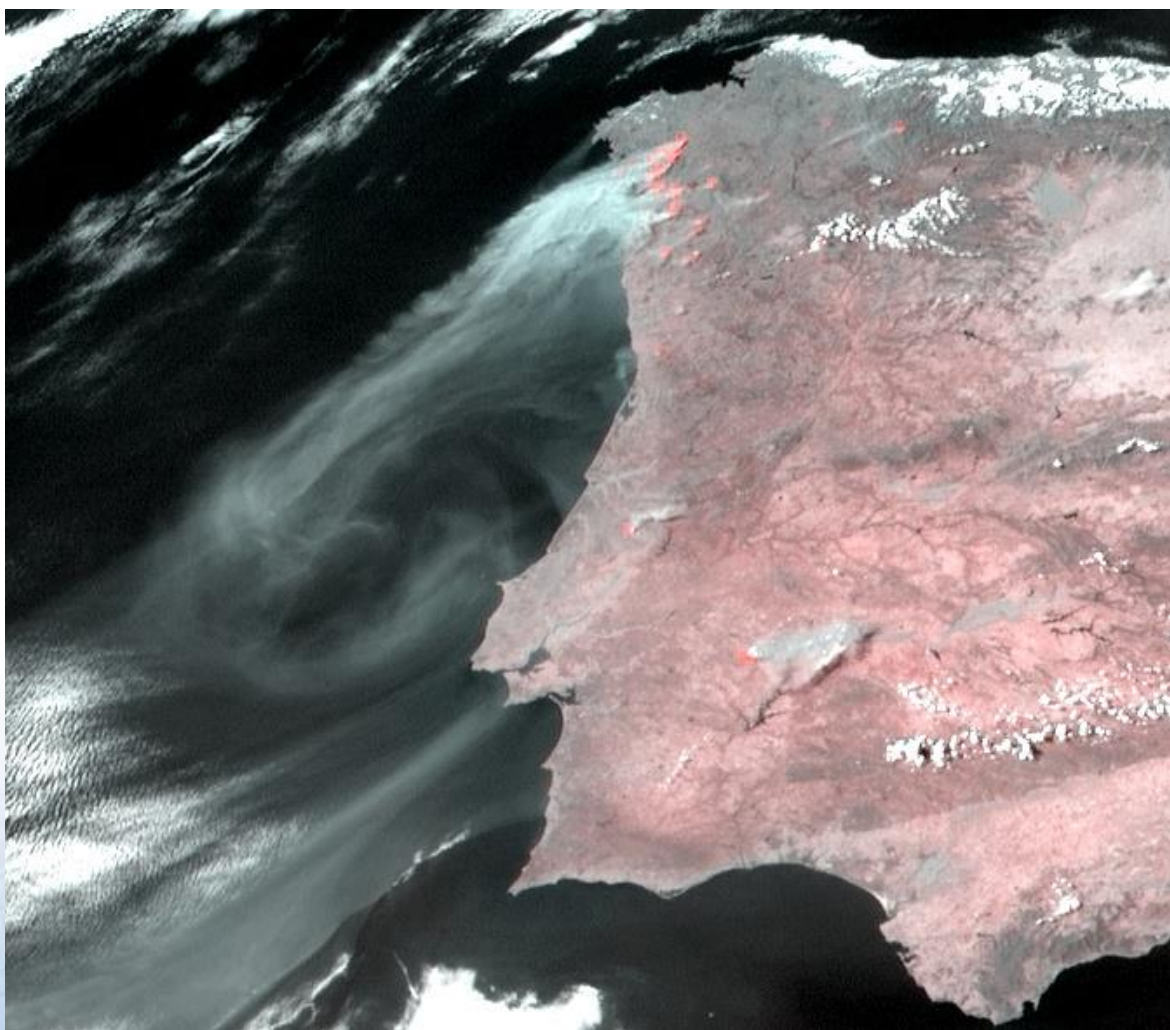
Physical concepts

- Sub-pixels effects

Remote sensing of fires and vegetation

- Scars in solar channels
- Land products
- Smoke

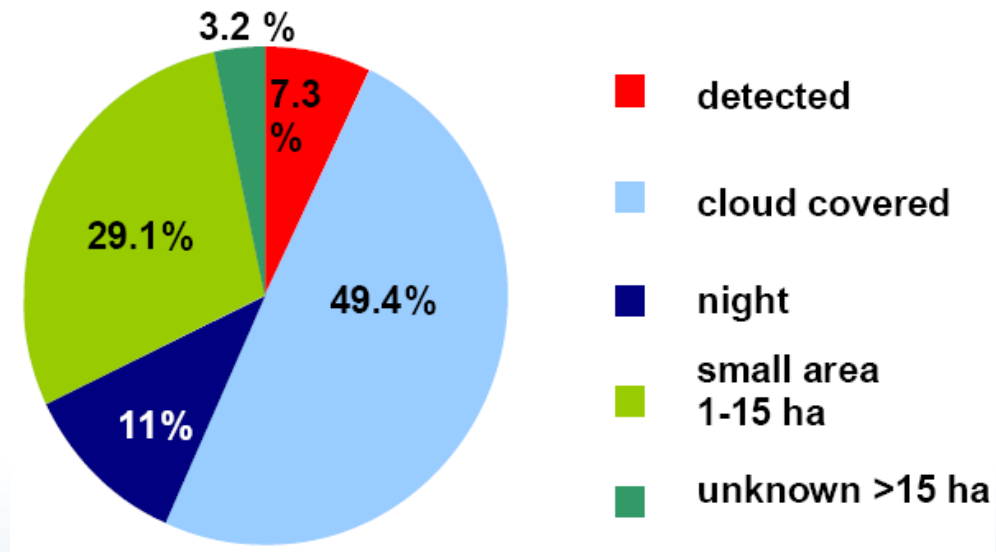
Exercise and quiz



Meteosat: 30% of channel 3.9 μ m on top of HRV

2006-July-7 16:00

Satellite detection of fires



- Study in Croatia, N Strelec, for 2007-2009
- Average: 750 fires/year, 160.000 Ha , $\langle x \rangle = 1.5$ km average horizontal dimension
- Poor performance of satellites due to: small active surface, cloud or smoke cover, location uncertainty (for fire brigades)

SEVIRI CHANNELS

		Properties			
Channel	Cloud	Gases	Application		
HRV 0.7	Scattering 0 -----> Emissivity -----> Absorption -----< 1	Broad band VIS	Surface, aerosol, cloud detail (1 km)		12
VIS 0.6		Narrow band	Ice or snow		1
VIS 0.8		Narrow band	Vegetation		2
NIR 1.6		Window	Aerosols, snow<>cloud		3
IR 3.8		Triple window	SST, fog<> surface, ice cloud		4
WV 6.2		Water vapour	Upper troposphere 300 Hpa humidity		5
WV 7.3		Water vapour	Mid -troposphere 600 Hpa humidity		6
IR 8.7		Almost window	Water vapour in boundary layer, ice<>liquid		7
IR 9.7		Ozone	Stratospheric winds		8
IR 10.8		Split window	CTH, cloud analysis, PW		9
IR 12.0	Split window	Land and SST		10	
IR 13.4		Carbon dioxide	+10.8: Semitransparent-cloud top , air mass analysis		11

Why 3.9 μm for fire?



Please type short suggestions here

3.9 μ m characteristics: mark the true!

Maximum emission by flames

Response to subpixel thermal anomalies

Small sun contribution

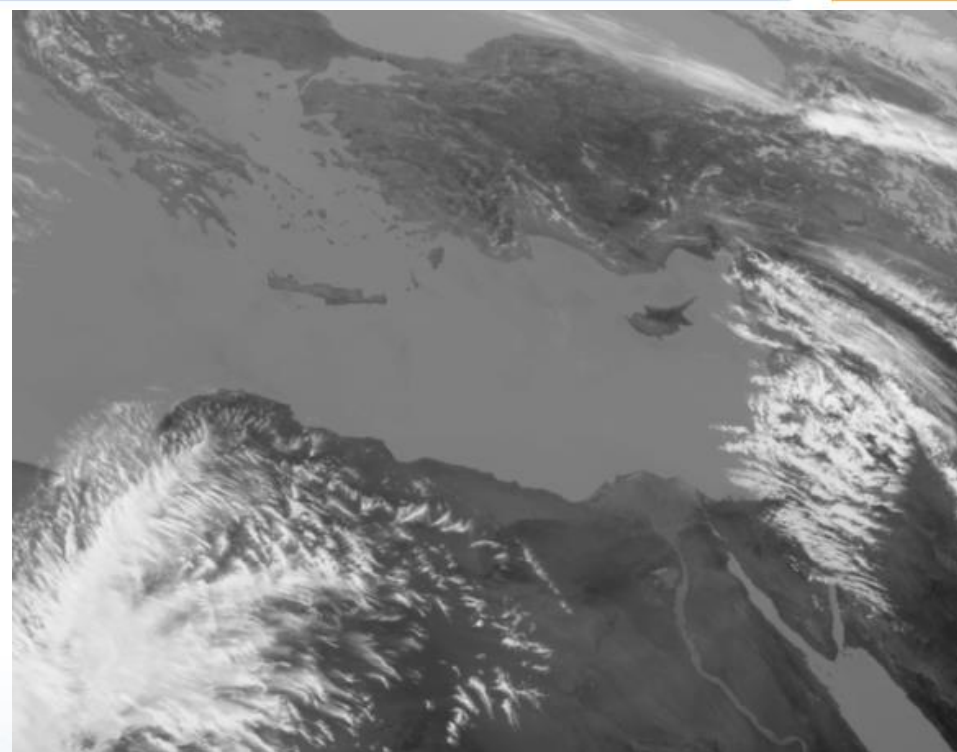
No absorption by water vapour

No absorption by carbon dioxide

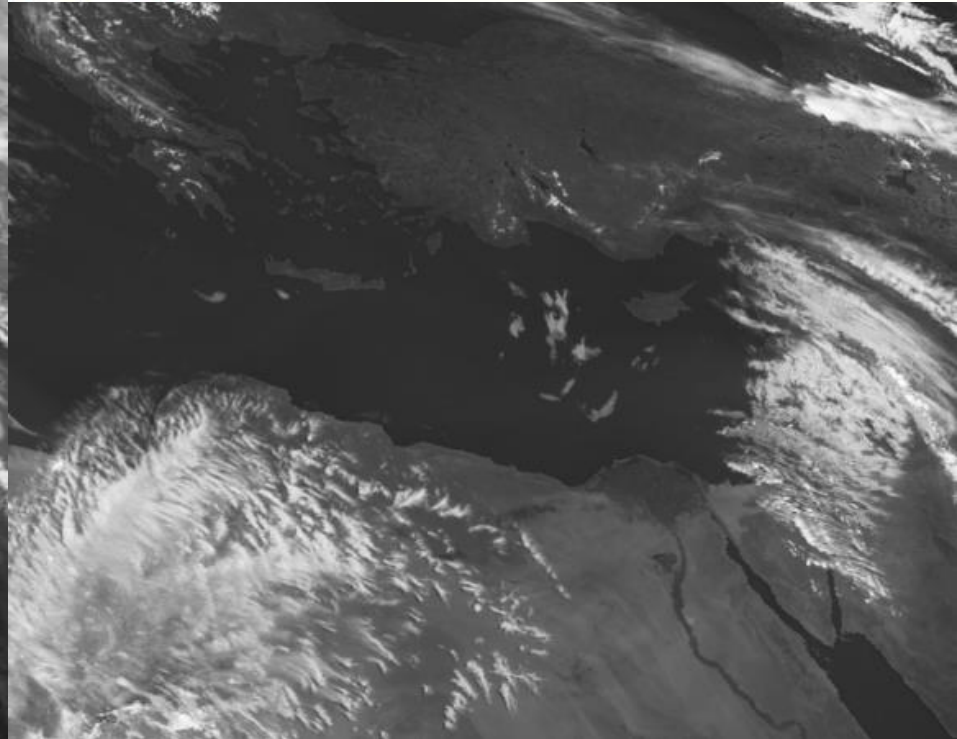
Meteosat pixel saturation for fires

Low ground emissivity

3.9 μm and 10.8 μm channels: IR window channels



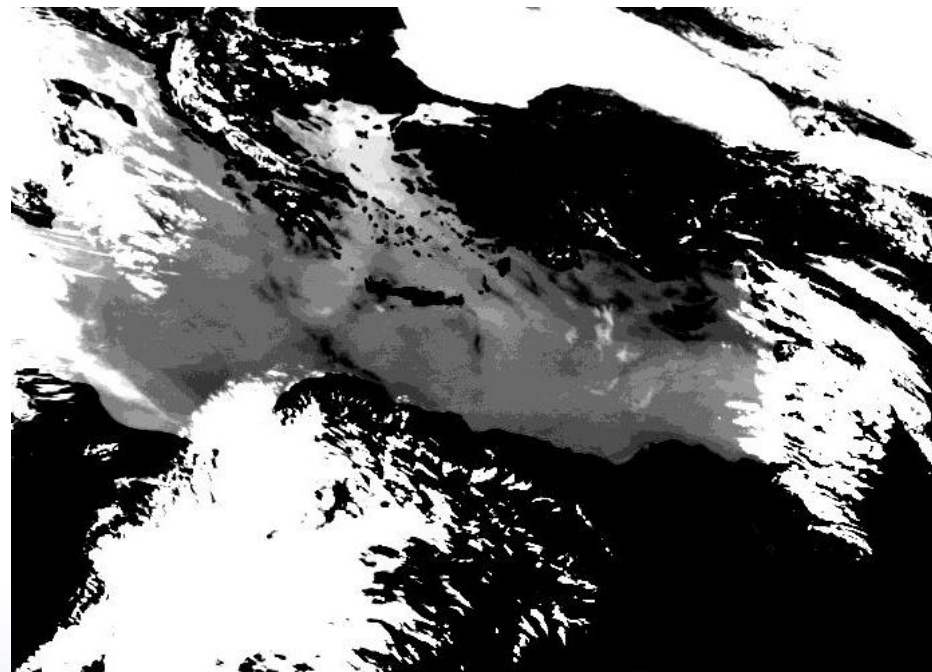
2016_Apr_05 10UTC Channel 10.8 μm [215K .. 315 K]



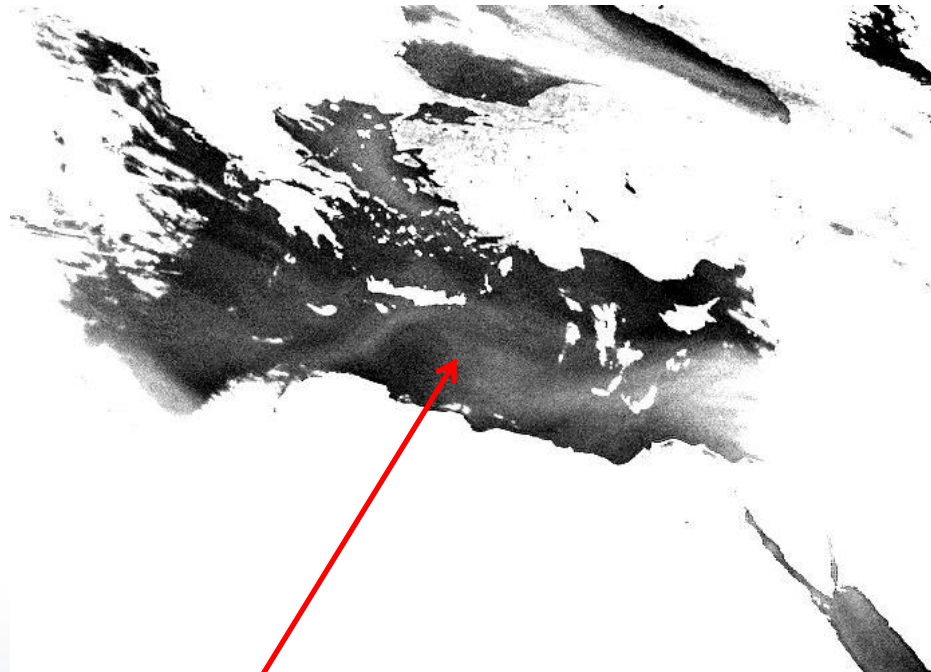
Difference 3.9 μm - 10.8 μm [-8K .. +60K]

Differences 3.9 μm – 10.8 μm due to:

Sun (+20K), gas absorption (-5K), ground type(+/-3K) and ... Planck (+5K)



2016_Apr_05 10UTC Channel 10.8 μ m [215K .. 315 K]



Difference 3.9 μ m - 10.8 μ m [-8K .. +60K]

Over water, 10.8 μ m roughly shows SST fields
But 3.9 μ m – 10.8 μ m shows humidity at low level

3.9 μm and 10.8 μm : window channels

3.9 μm

- ❖ Negligible absorption by atmospheric humidity
- ❖ Close to a **CO2 absorption** band, 4-7 Kelvin signal reduction
- ❖ High temperature **sensitivity** (big sub-pixel effects) $\sim 14 * \Delta T/T$
- ❖ **Blinding** effect by hot pixels, affecting measurements west of the saturated pixel
- ❖ Fog warnings: daytime start or night dissipation onset
- ❖ **Sun** enhancement during day, emission effects during night

10.8 μm

- ❖ 1-2 Kelvin absorption by atmospheric humidity
- ❖ No signal reduction by CO2
- ❖ Lower temperature sensitivity (small subpixel effects) $\sim 4 * \Delta T/T$
- ❖ No risk of sensor blinding by fires
- ❖ Low values compared with 3.9 μm due to semitransparent cloud or smoke.

T-Difference 4-9

1

-5K

-2K

4

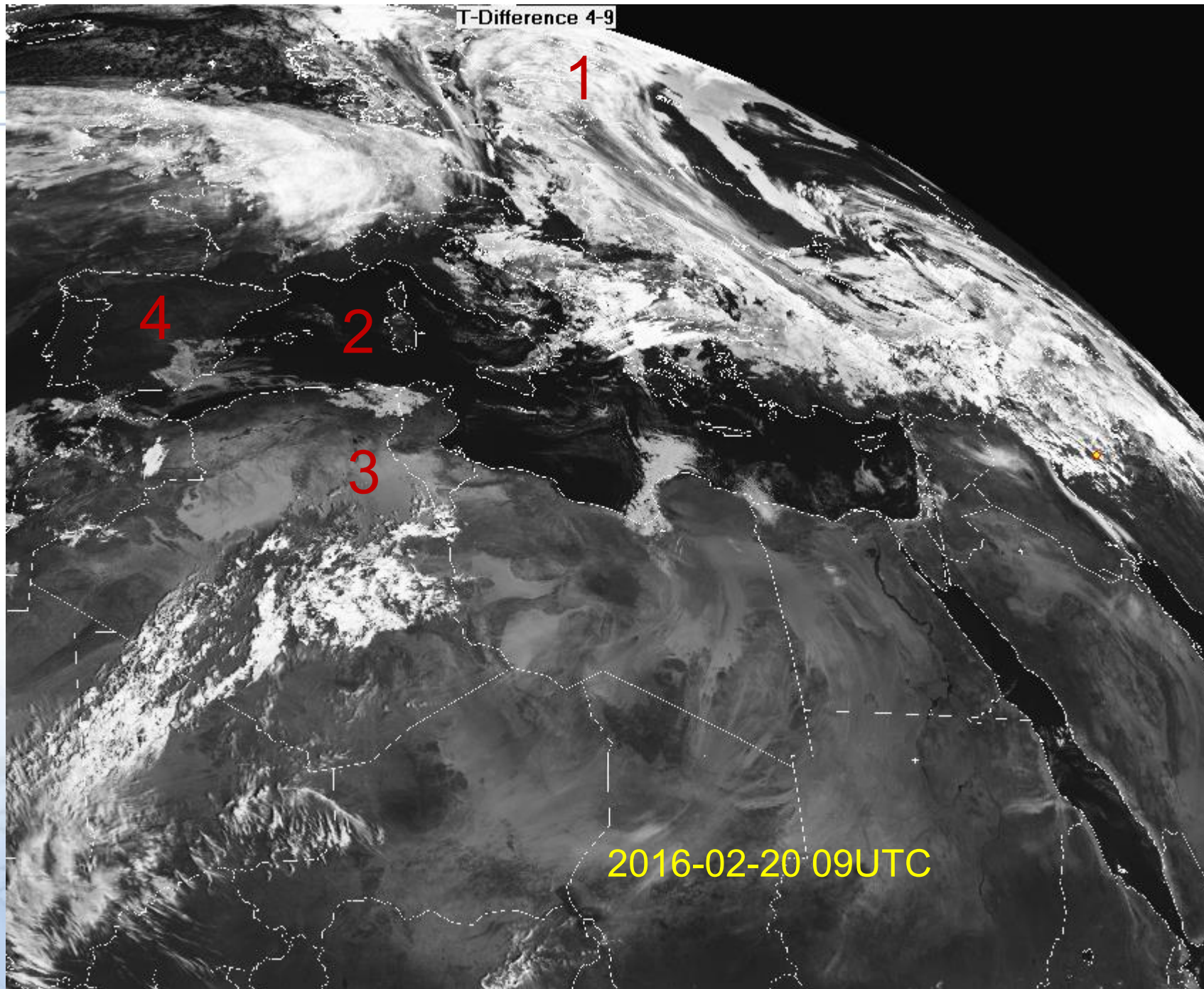
2

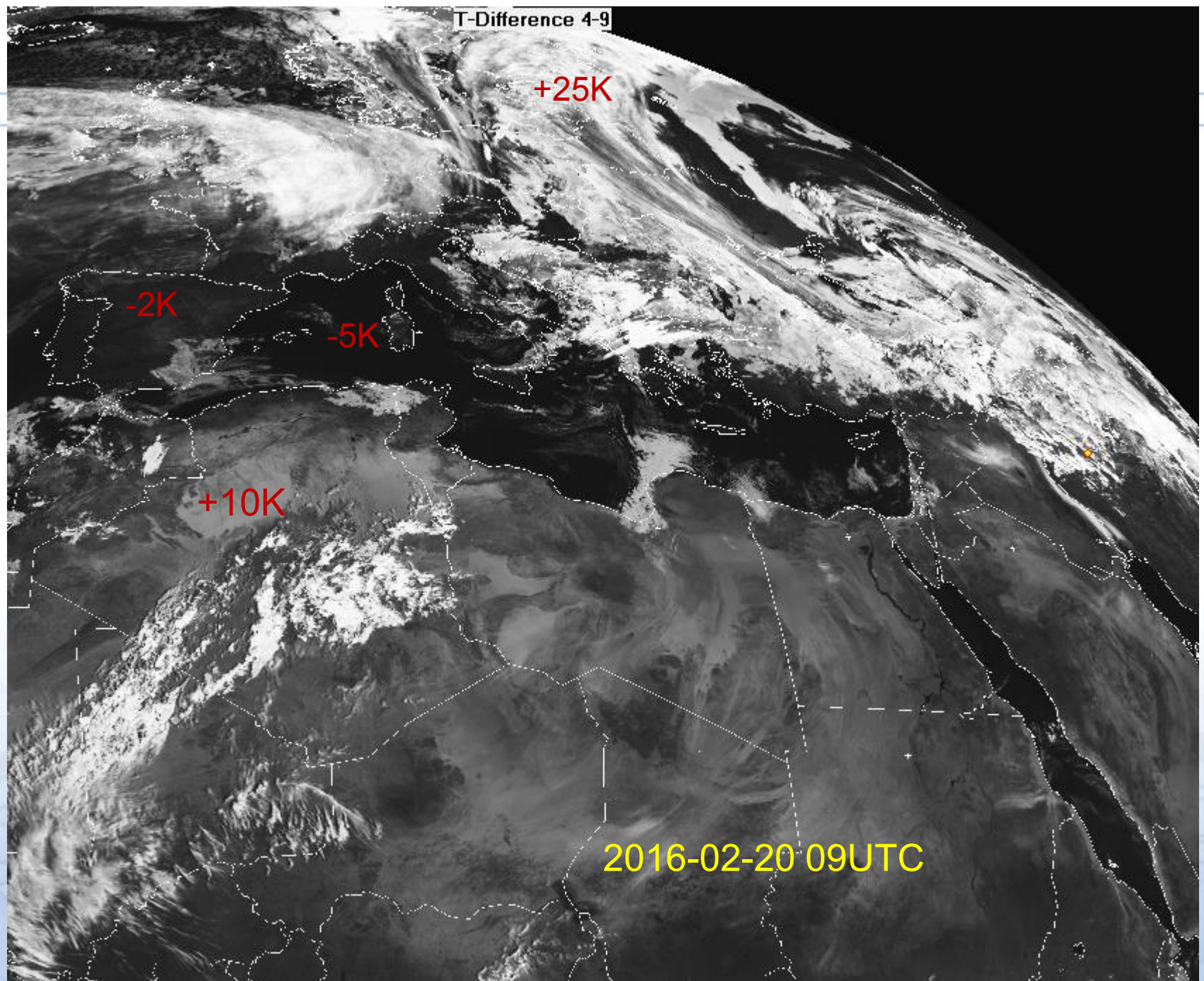
+10K

3

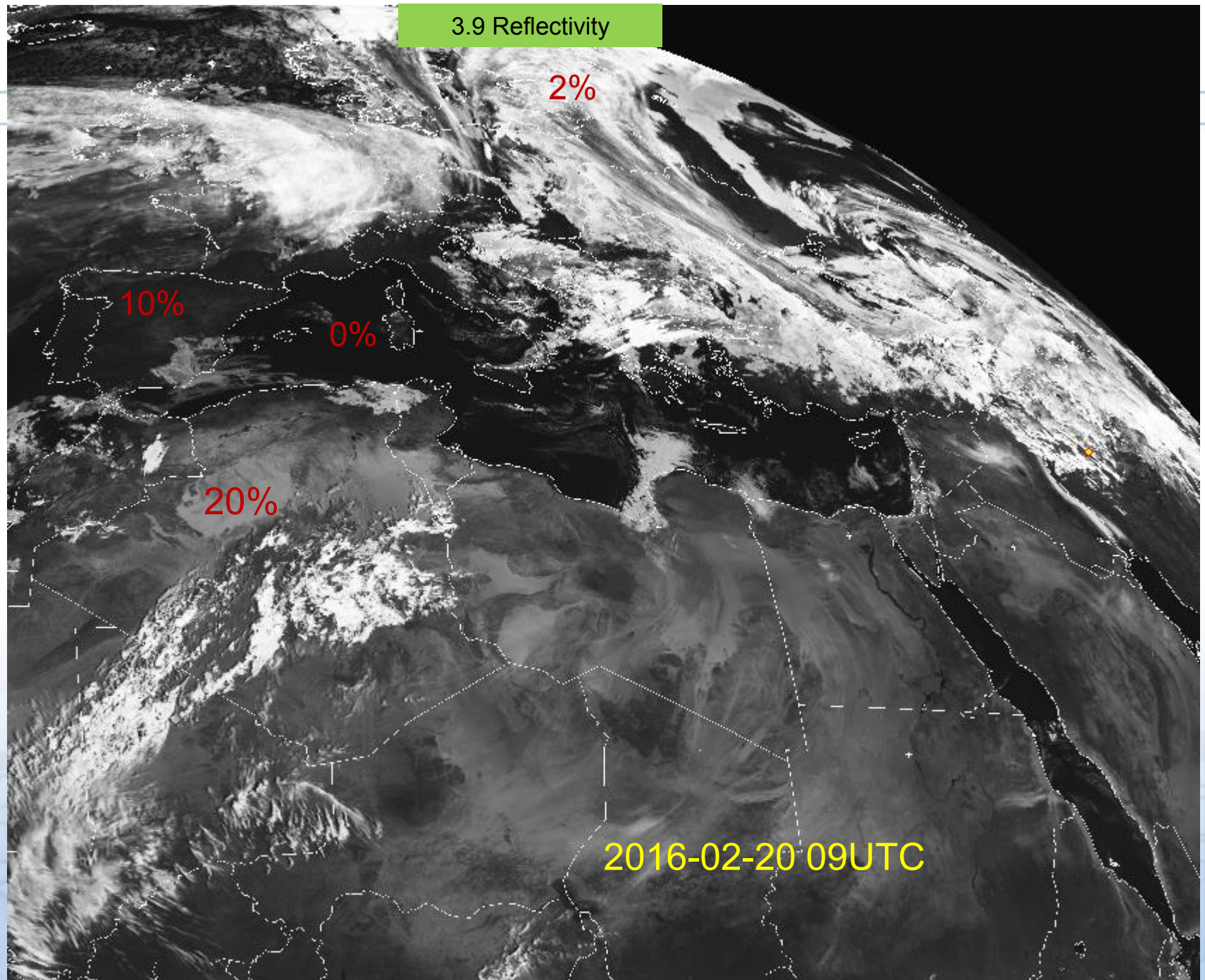
+25K

2016-02-20 09UTC

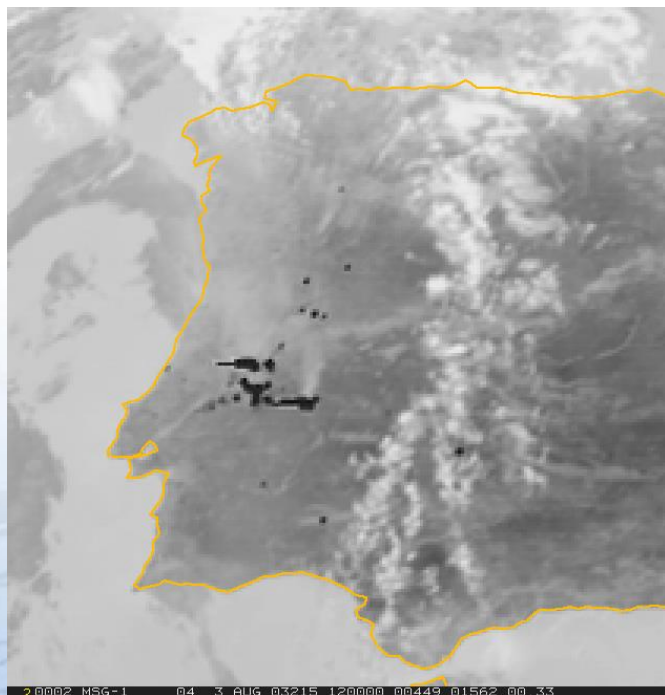
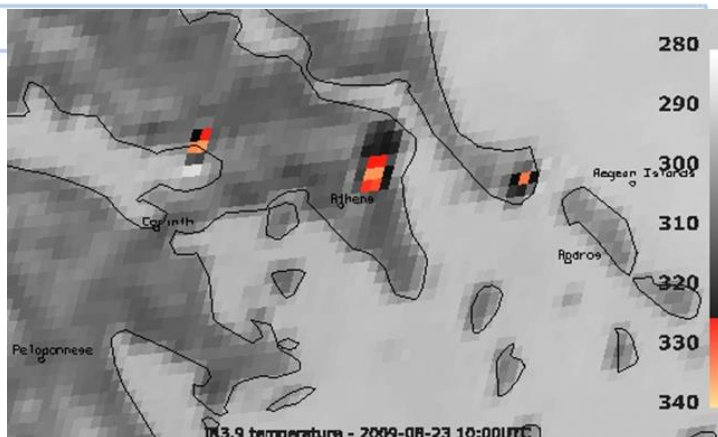




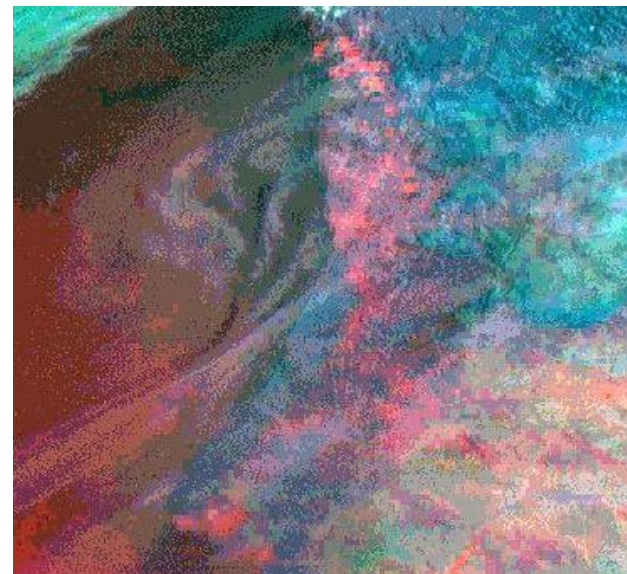
3.9 Reflectivity



3.9 μm and 10.8 μm channels: sensor blinding and filters

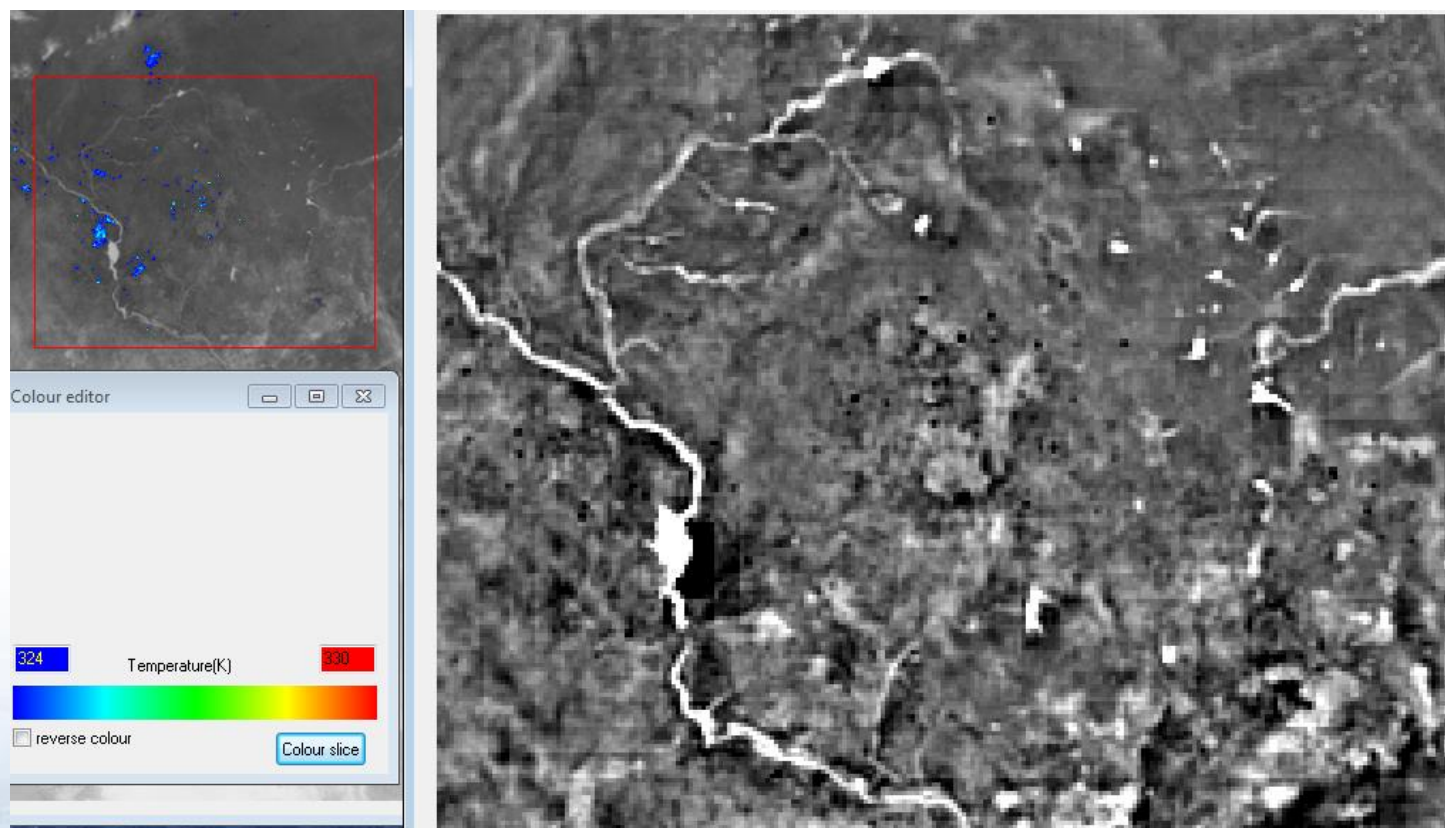


For pixels west of the fire the sensors can be blinded, following geometrical patterns (**rings**)



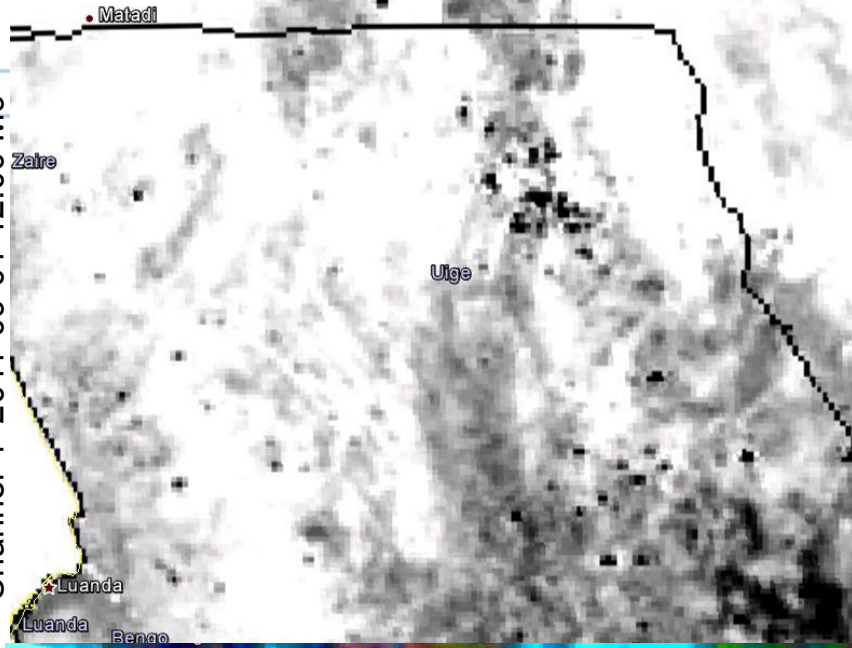
2006_08_07 06-19UTC
rgb_12 + 4-3-2
HRV can be combined with
lower horizontal resolution for
more spectral information



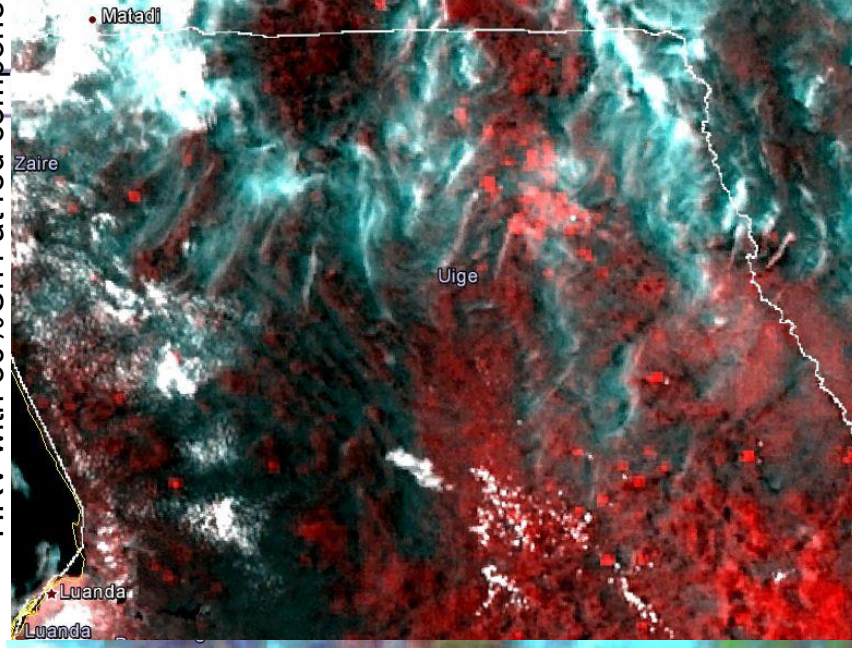


2015-02-09_15UTC, channel 4, Meteosat-10
Contrast with neighbours is better than thresholds in 3.9 μm to spot fires

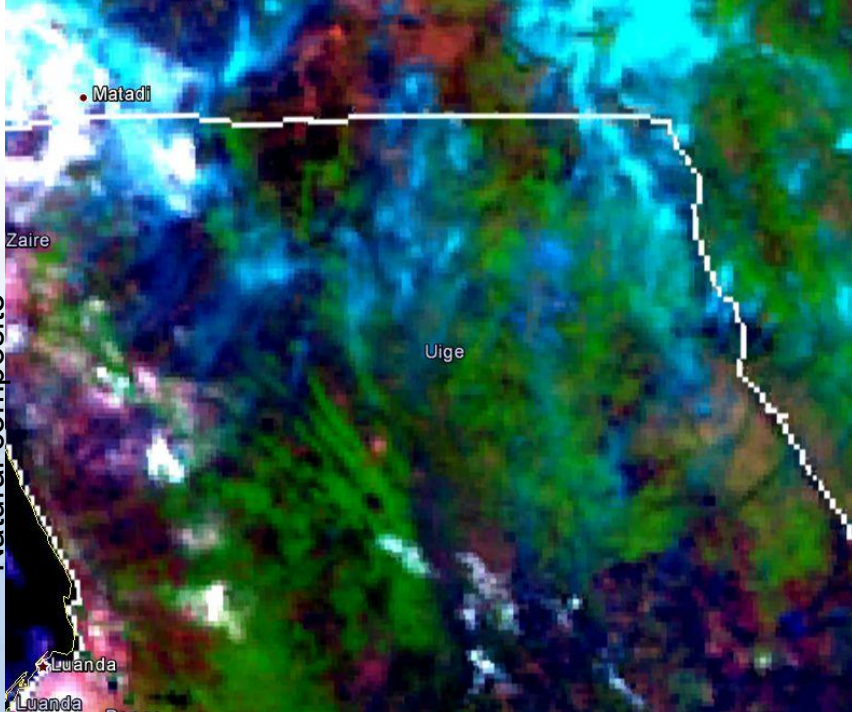
Channel 4 2011-09-01 12:00 M9



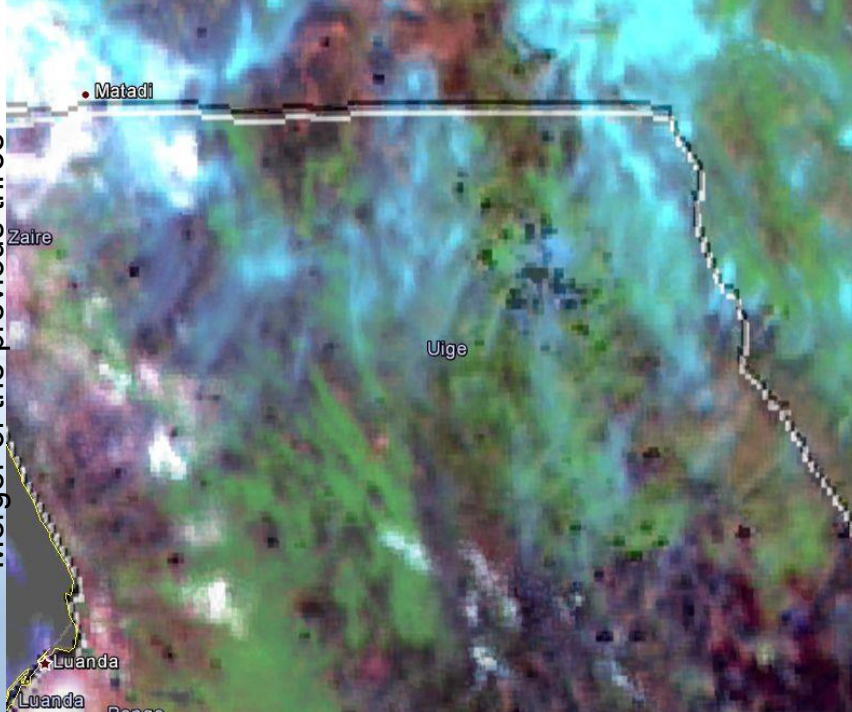
HRV with 30%Ch4 at red component



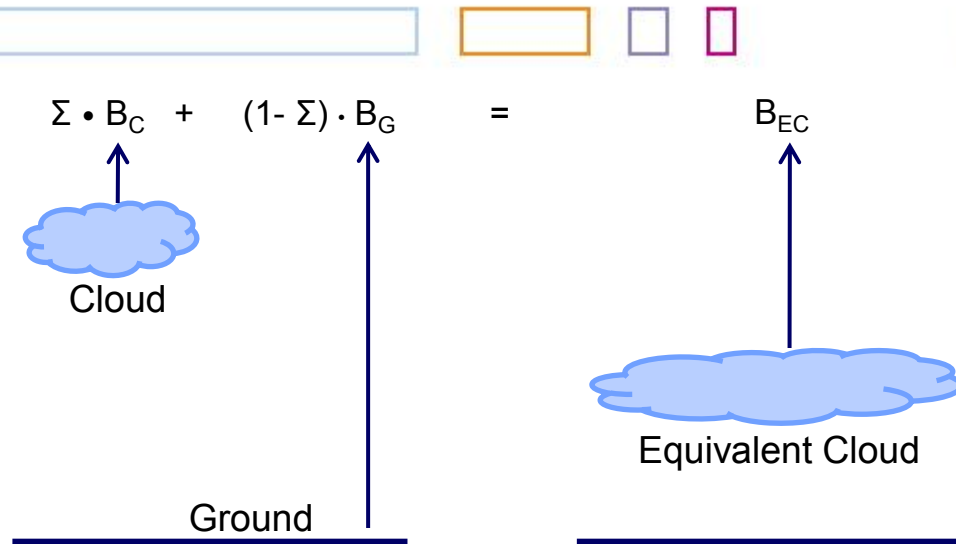
Natural composite



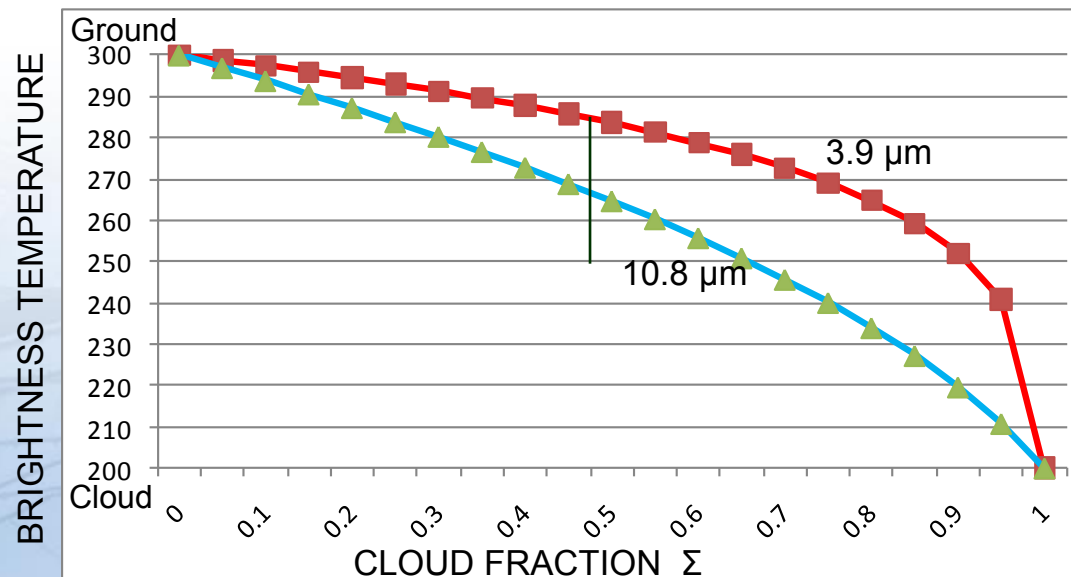
Merger of the previous three



Sub-pixel effects = temperature sensitivity = warm bias



The equivalent temperature is not the average temperature, but shows a WARM BIAS!



50% cloud gives:

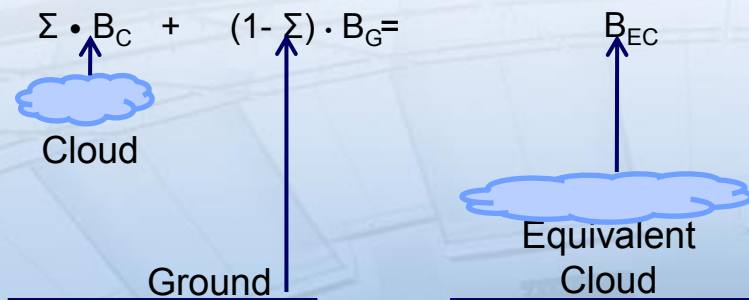
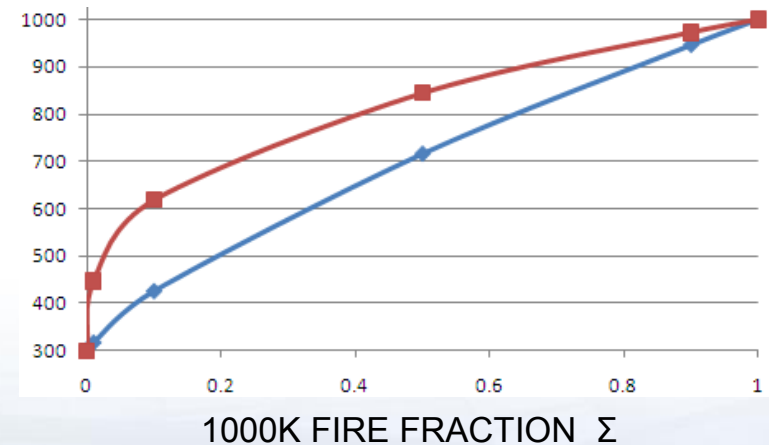
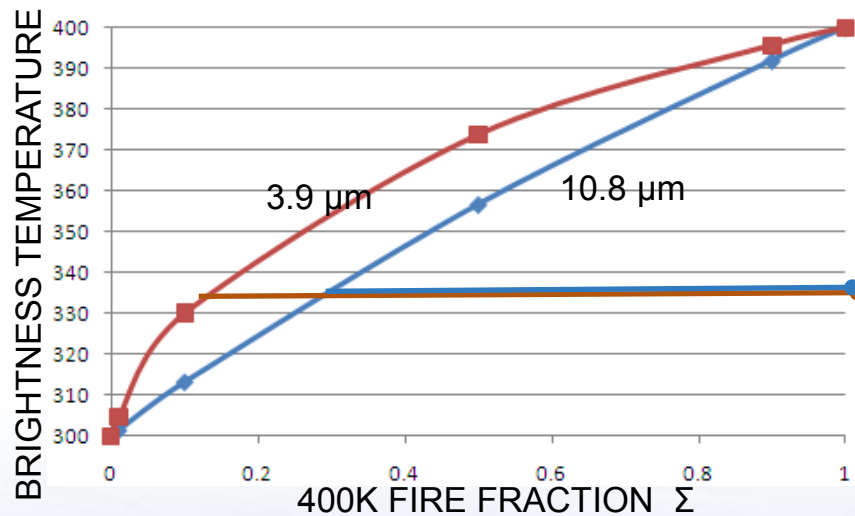
282 K (at 3.9 μm)

264 K (at 10.8 μm)

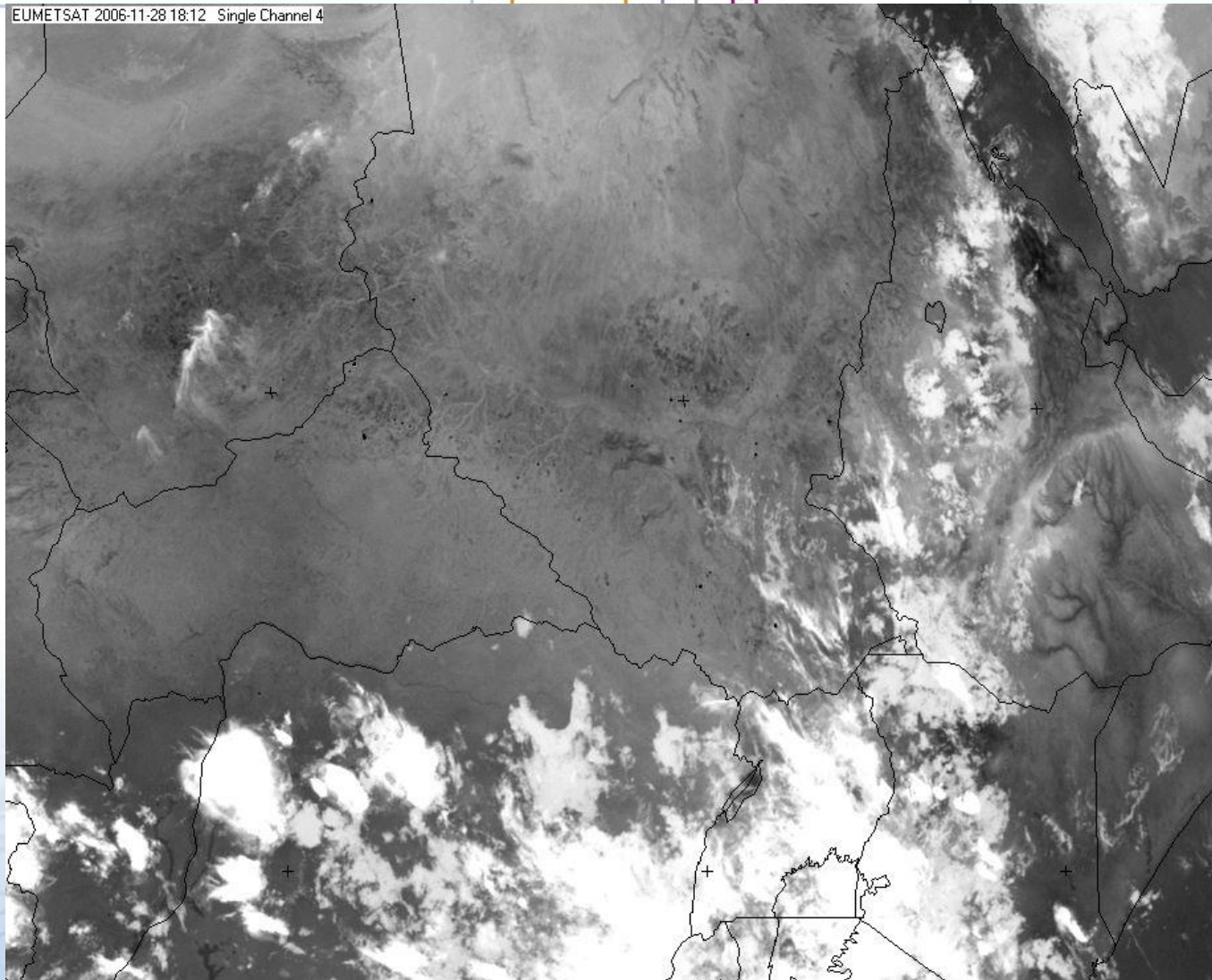
and NOT the average 250 K

Subpixel effects = temperature sensitivity = warm bias

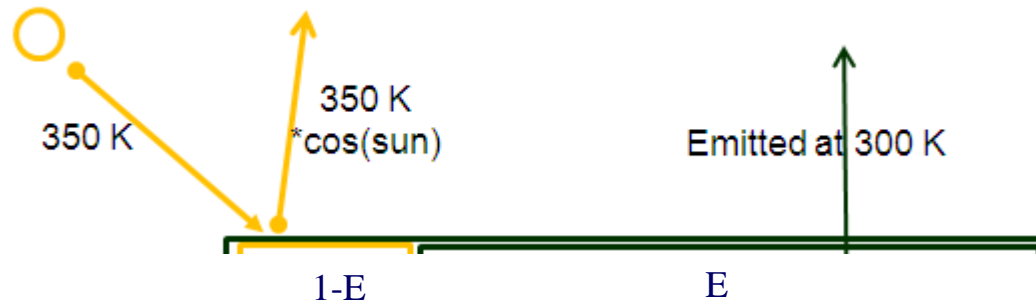
Widespread fires (15%) show less difference 3.9 μ m – 10.8 μ m than small ones (5% of the pixel)



Fires on 1.6μm images



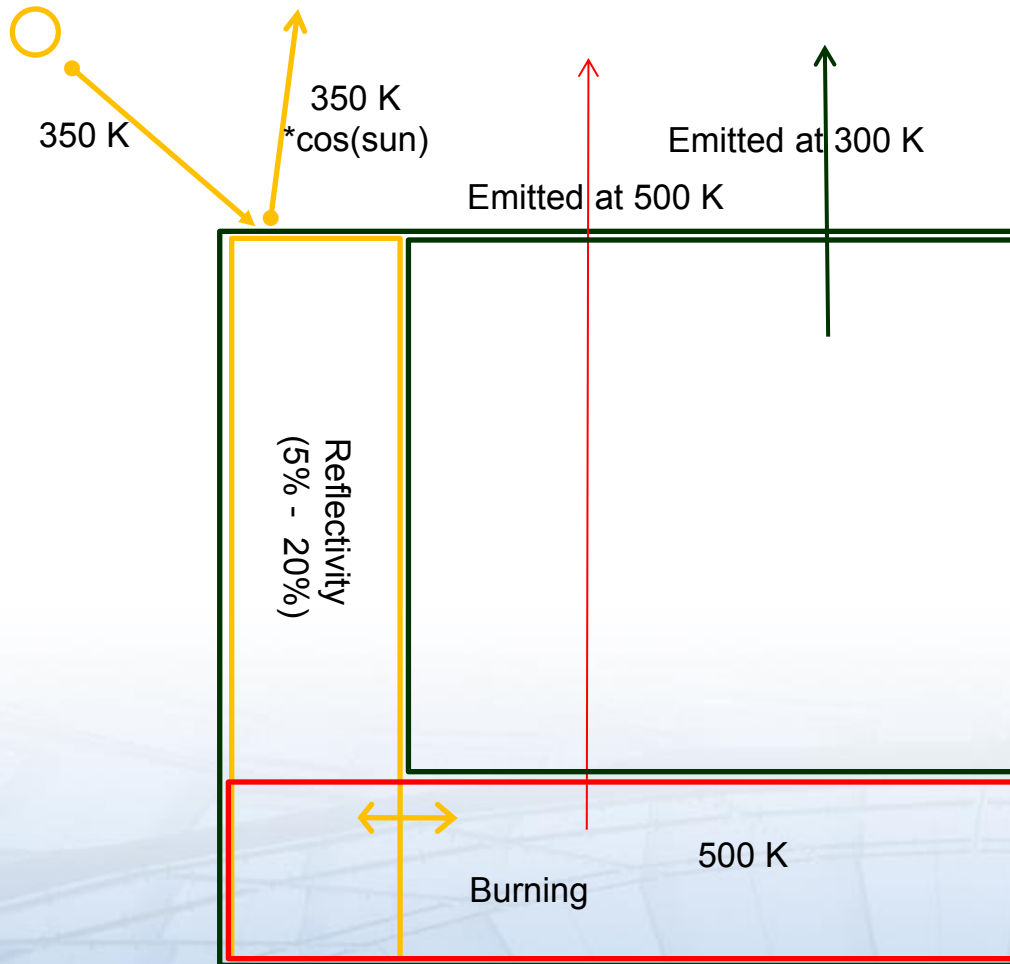
Solar reflection and emission together (3.9 μm)



$$B(BT) = (1-E) * B(350K) + E * B(300K)$$

- Warm bias in brightness temperature towards 350K (depends on illumination)
- During night, brightness temperature (BT) is lower than 300K
- Albedo (1-E) varies with type of soil: 20% (savannah) to 5% (forest)
- Cloud (1-E=2%) is usually present in burning areas

Hot spots contributions in a pixel (3.9μm)



		Reflectivity 3.9μm	
DAY BT		5%	20%
Fraction burning	0	314	333
	0.01	328	339
	0.1	380	370
	0.5	449	425
	1	490	460
NIGHT BT		5%	20%
Fraction burning	0	296	284
	0.01	318	304
	0.1	377	356
	0.5	448	421
	1	489	457

Sunrise and sunset change 3.9μm BT but normally **outside of the detection** range of SEVIRI

Not only 3.9µm allows fire detection

NEAR INFRARED (e.g. 1.6µm)

- More adequate for smoke detection than 3.9µm
 - Small fires not visible (below threshold)
- No CO₂ absorption (higher “fire temperature”)
 - High sub-pixel sensitivity

Karthala, Met-8, 29 May 2006, 12:15 UTC
Natural colours RGB 1.6µm-0.8µm-0.6µm

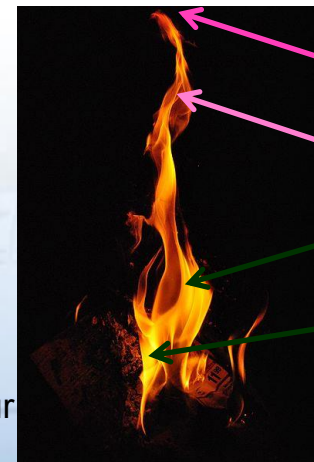


How hot is lava?

3.9µm

- Hotspots are easily detected
- Total absorption of ground radiation by CO₂
- BT is temperature of the **CO₂ layer** above the fire
 - 100m minimum fire size for Meteosat pixel
- Sun interference noticeable (~20 K), but truncated by 3.9µm channel dynamic range limit (333K)
- Difficult statistics due to man-made fire generation (e.g. after harvest)

Abundant CO₂
..and soot...



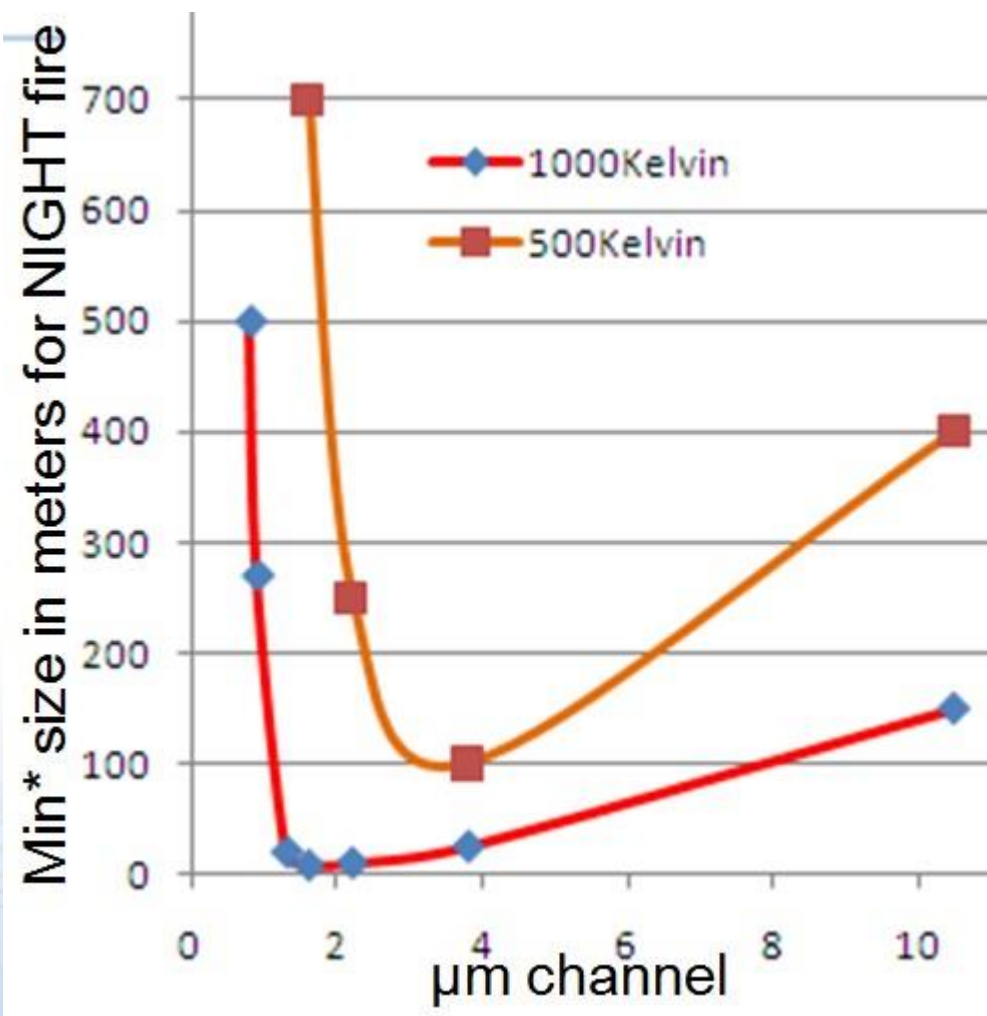
300K,
neighbour
pixels

Which one is the “fire
temperature”?

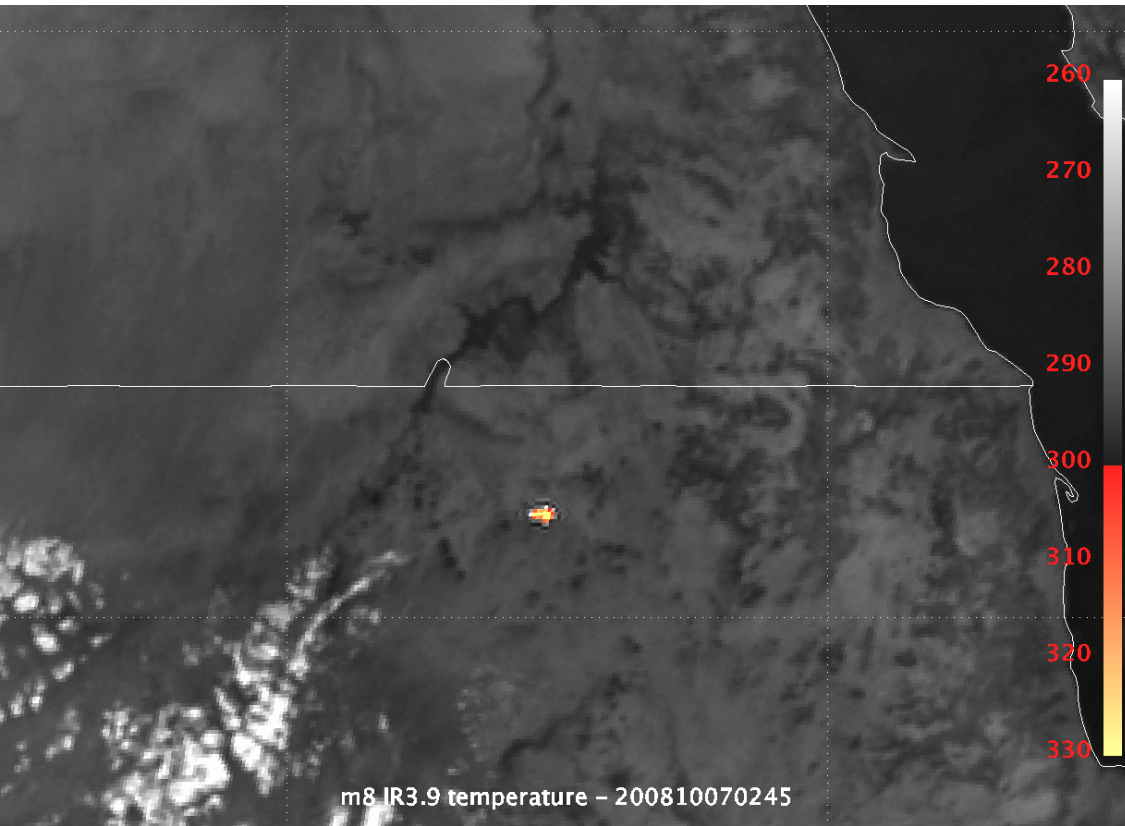
The fire traffic lights



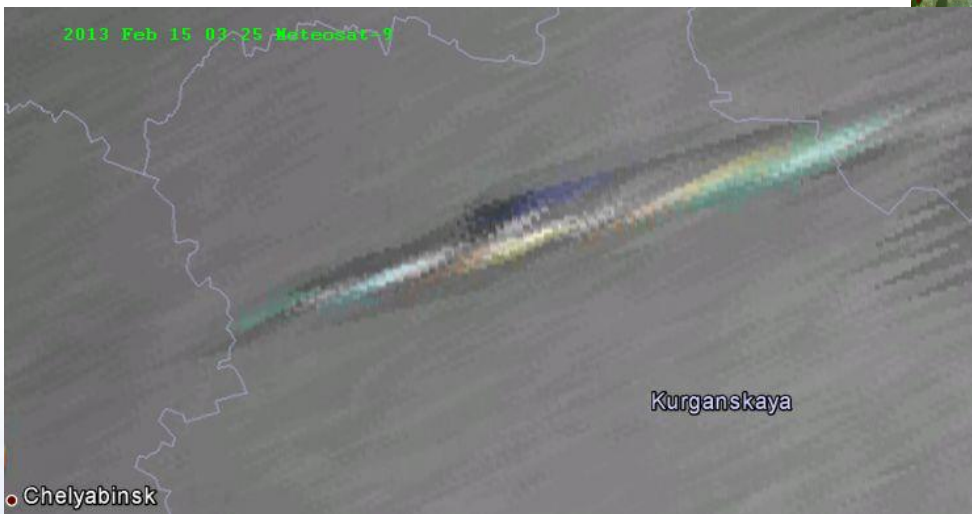
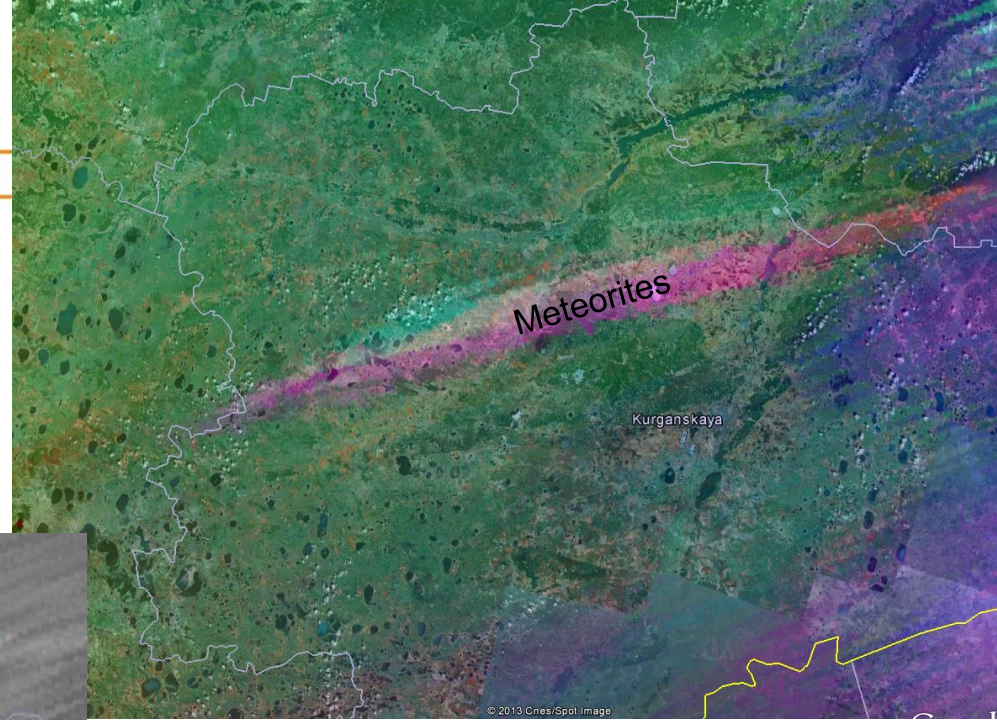
- A fire at 500K will be sensed, as it grows
- **first** by 3.9 μm (at ~100m)
- **second** by 2.2 μm (250m)
- **third** by 10.8 μm (400m)
- An RGB=(3.9;2.2;10.8) might be a good indicator for severity of a fire.
- For a hotter fire (1000K), typically gas flares, channels in the solar domain react faster than 3.9 μm



Meteorites on 3.9 μm images



Subpixel detection at 3.9 μm

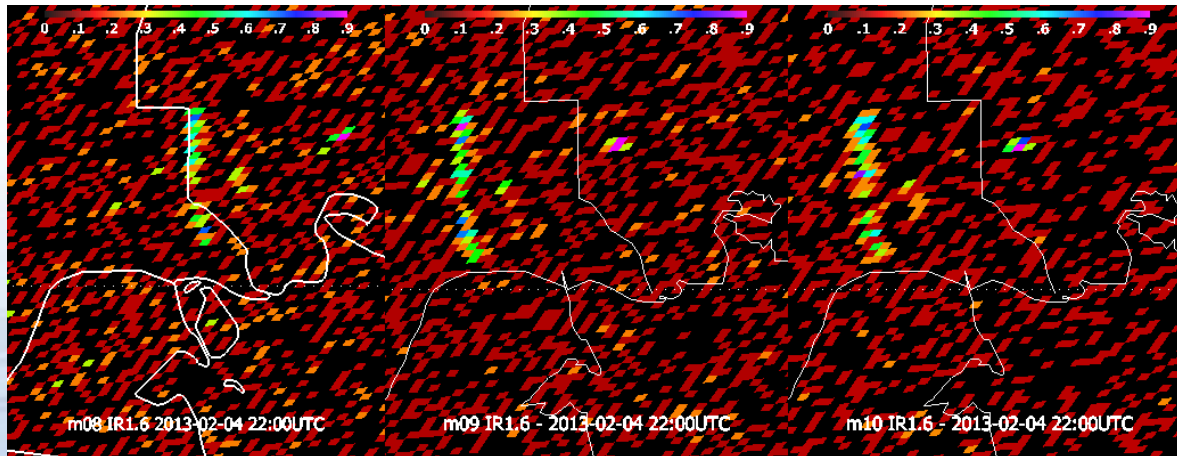


Colour from Meteosat-9 channel 3.9 μm .
Blue=270K Red=280K

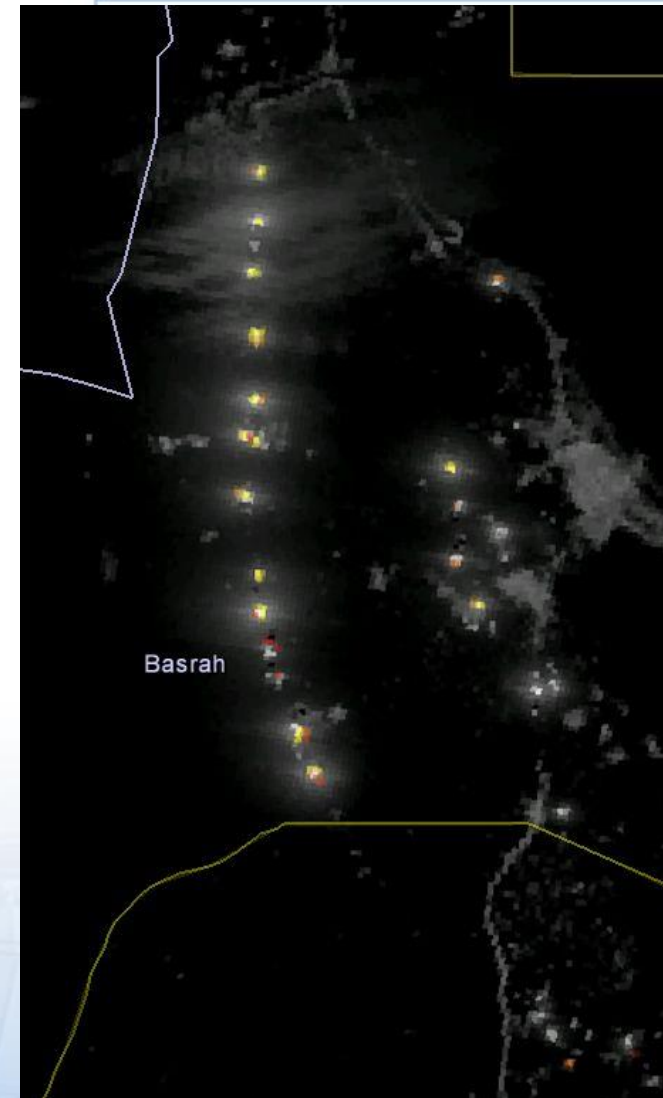


Meteosat IR dynamic range top limits (kelvin)

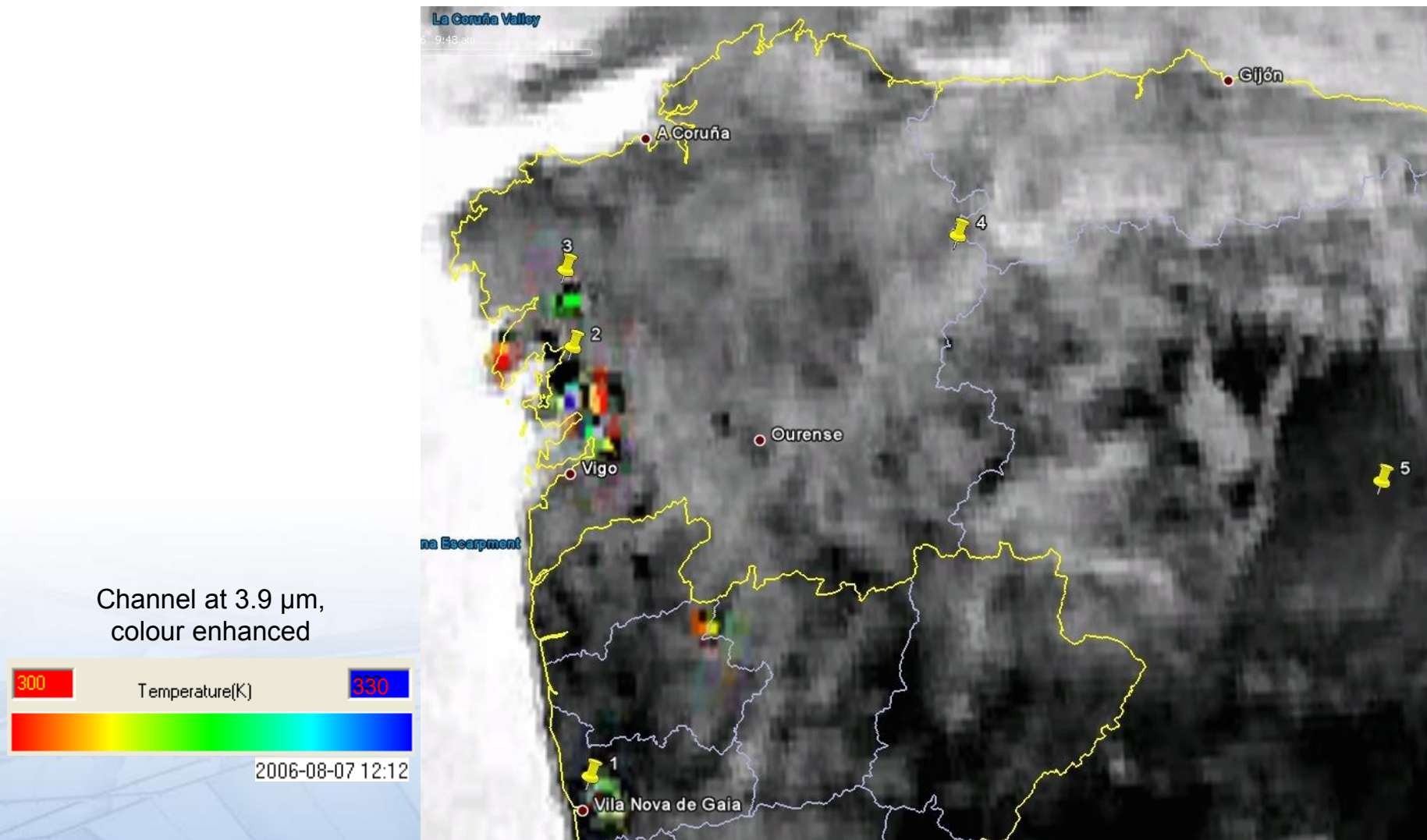
Channel (μm)	3.8	8.7	9.7	10.5	12.3	13.3
Absorber	CO ₂	Sx	O ₃	small	H ₂ O	CO ₂
Dynamic MSG	335	300	310	335	335	300
Dynamic MTG	580	330	310	340	340	300



Meteosat-8,9,10 looking concurrently at gas flares in Kuwait through channel 1.6 μm



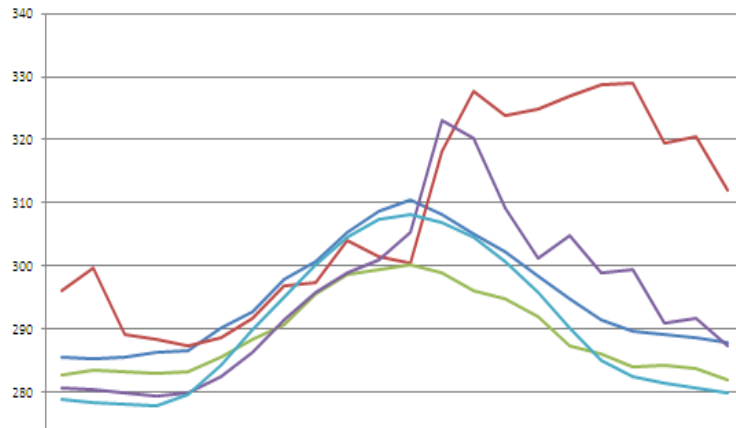
Fires in Galicia (Spain)



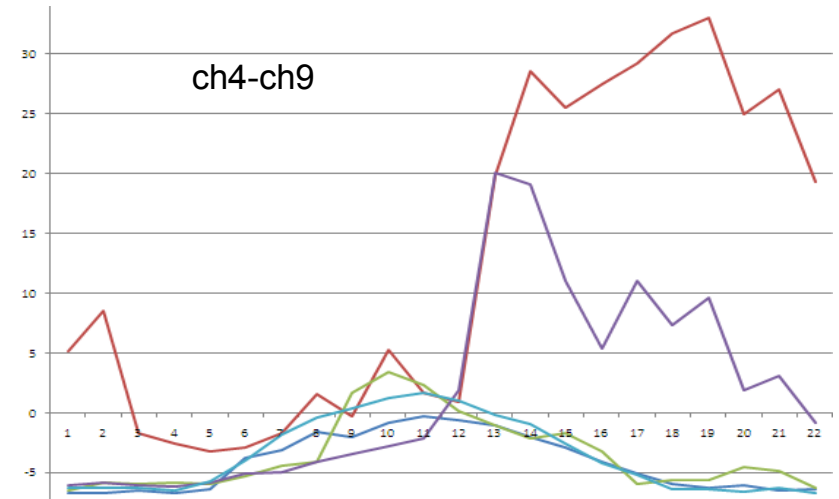
Hot spots, brightness temperature daily evolution



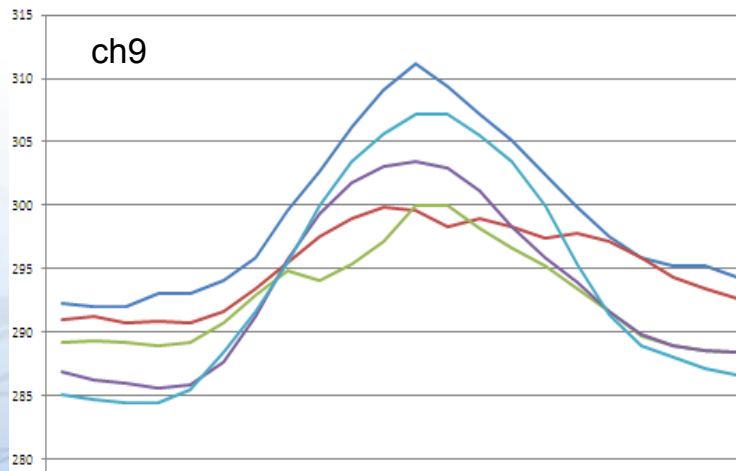
ch4



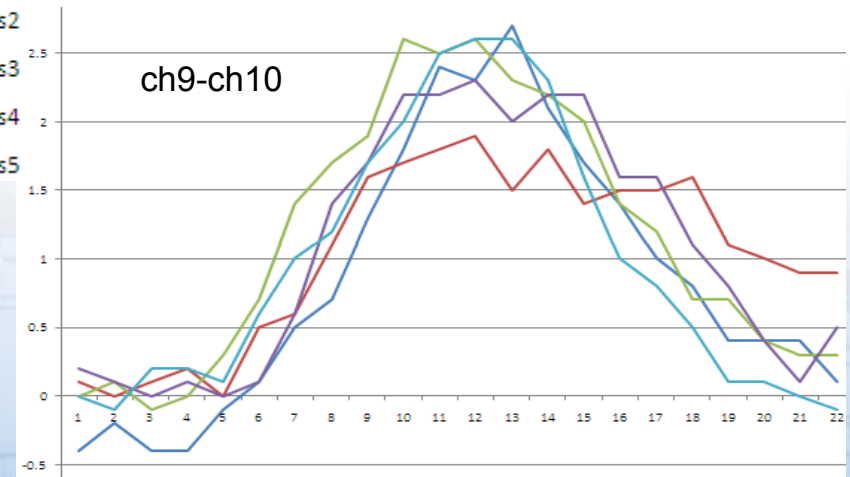
ch4-ch9



ch9



ch9-ch10



Series1

Series2

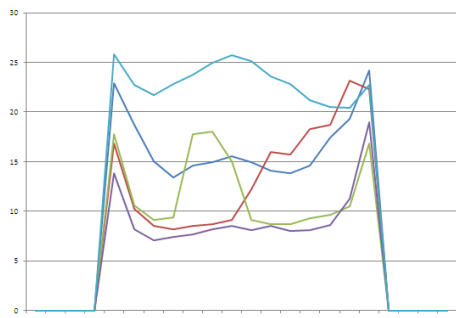
Series3

Series4

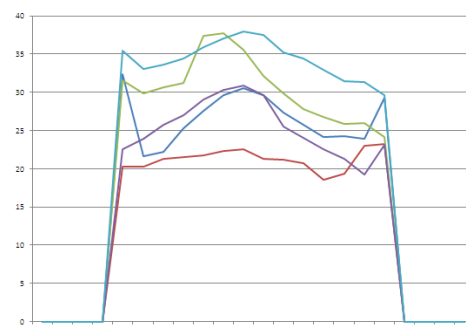
Series5

- Stronger response in 3.9 μ m than in 10.8 μ m or 12 μ m
- Optimal index is 3.9 μ m – 10.8 μ m
- Alternative index 10.8 μ m – 12 μ m, due to humidity increase?

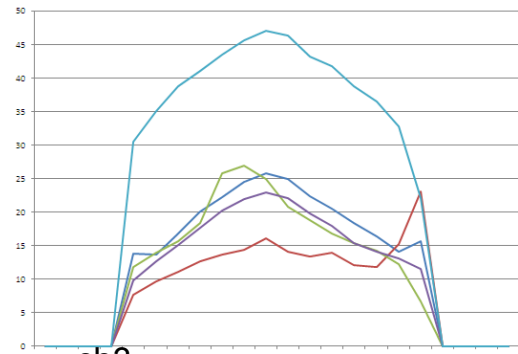
Solar reflection



ch1

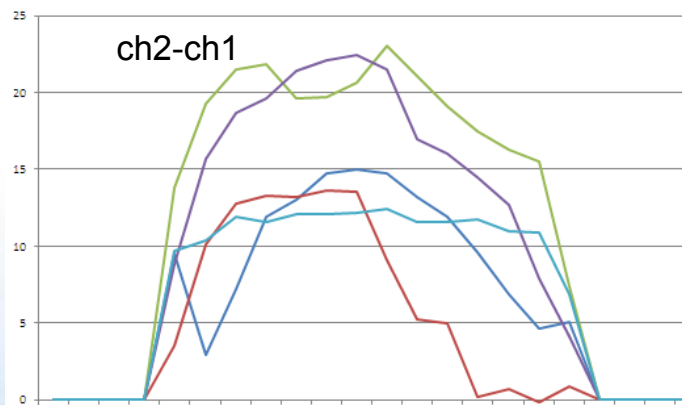
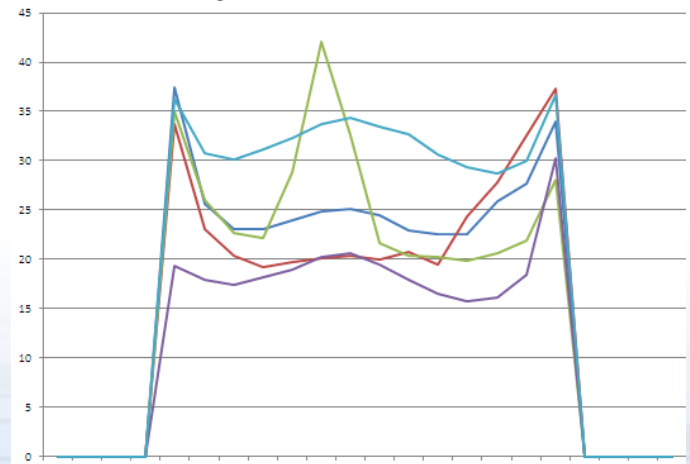


ch2



ch3

ch12



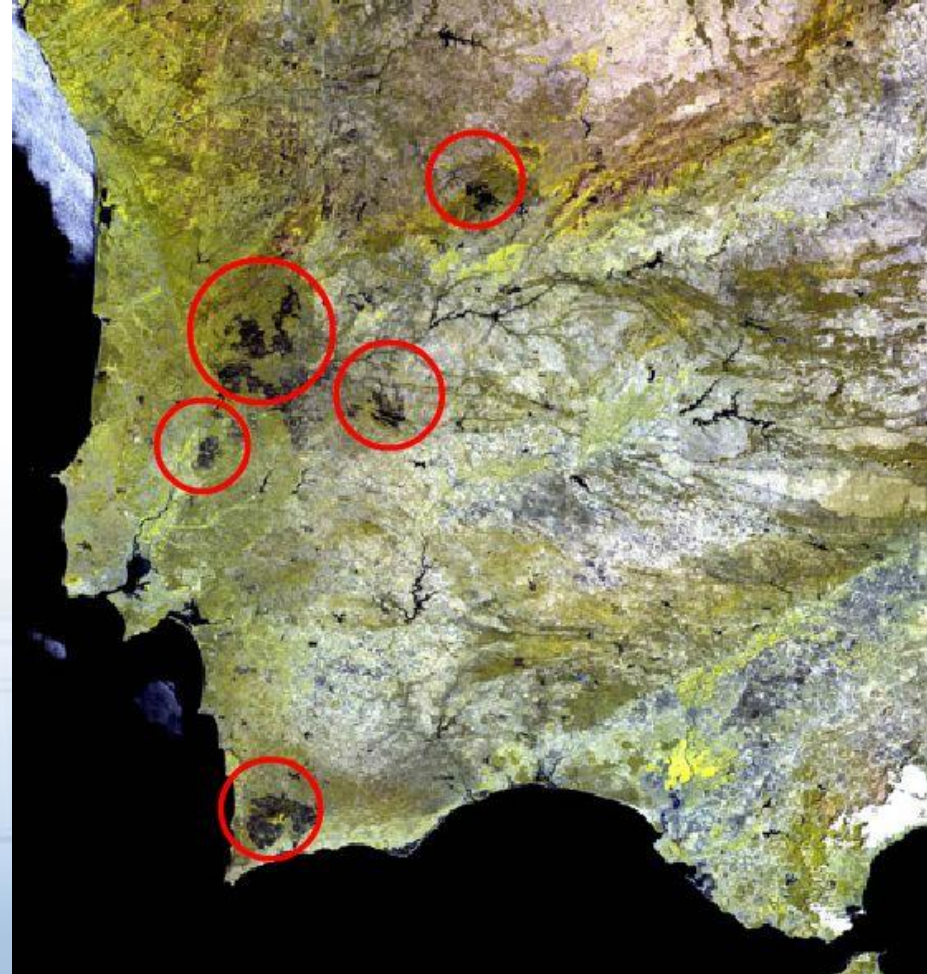
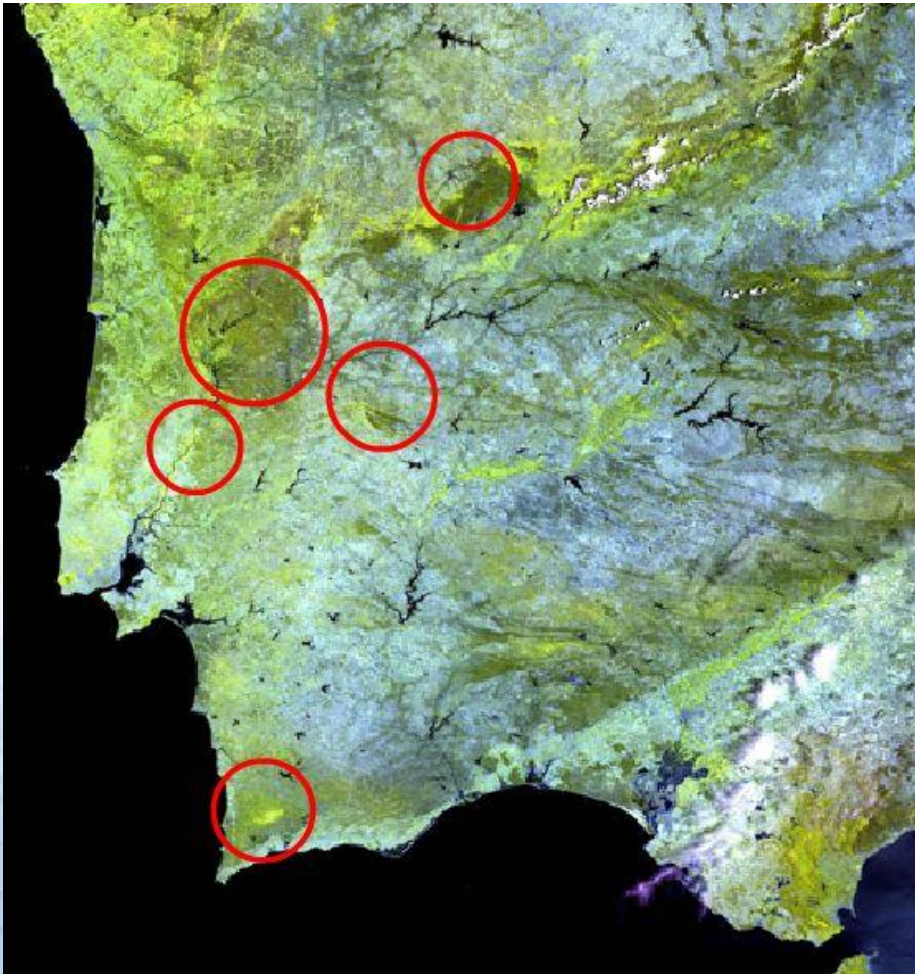
ch2-ch1

Series1
Series2
Series3
Series4
Series5

- $0.6\mu\text{m}$ reflection increases after the forest fire!
- More moderately for for $0.8\mu\text{m}$ and $1.6\mu\text{m}$

Effect of fire on vegetation can be measured by satellite

- Fires August 2003 Portugal: 5% of portuguese territory
- Scars can be evaluated on solar channels



LAND SURFACE ANALYSIS SATELLITE APPLICATIONS FACILITY



landsaf.meteo.pt

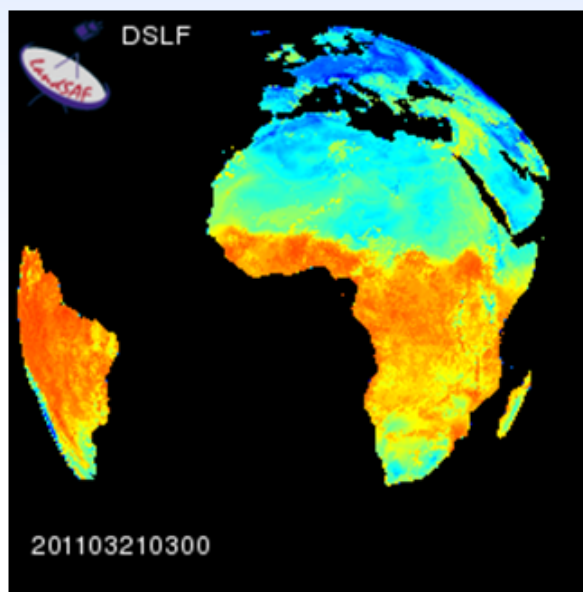
Home

The scope of Land Surface Analysis Satellite Applications Facility (LSA SAF) is to increase benefit from EUMETSAT Satellite (MSG and EPS) data related to:

- Land
- Land-Atmosphere interaction
- Biospheric Applications

The LSA SAF performs:

- R&D Programs.
- Operational Activities
- Generation
- Archiving
- Dissemination



[See colour legends...](#)

of land surface related products.

Latest News:

- **Important** IM Archive system maintenance. [see more...](#)
- **Important** IM Archive system maintenance. [see more...](#)
- **Information** LSA SAF Outage [see more...](#)
- **Information** LSA SAF Outage [see more...](#)
- **Update** MSG Images [see more...](#)

Product Development Status:

MSG/SEVIRI based products

Wild Fires

Fire Radiative Power - PIXEL

Fire Radiative Power - GRID

Vegetation Parameters

Fraction of Vegetation Cover

Leaf Area Index

Fraction of Absorbed Photosynthetic Active Radiation

Snow Cover

Snow Cover (daily)

Snow Cover (15 mins)

Other

Bi-Directional Reflectance Factor

Land Surface Emissivity

Albedo

Surface Albedo

MSG Ten Day Surface Albedo

Land Surface Temperature

Land Surface Temperature (15 mins)

Down-welling Surface Fluxes

Down-welling Surface Short-wave Radiation Flux

Down-welling Surface Long-wave Radiation Flux

Daily Downward Surface Shortwave Flux

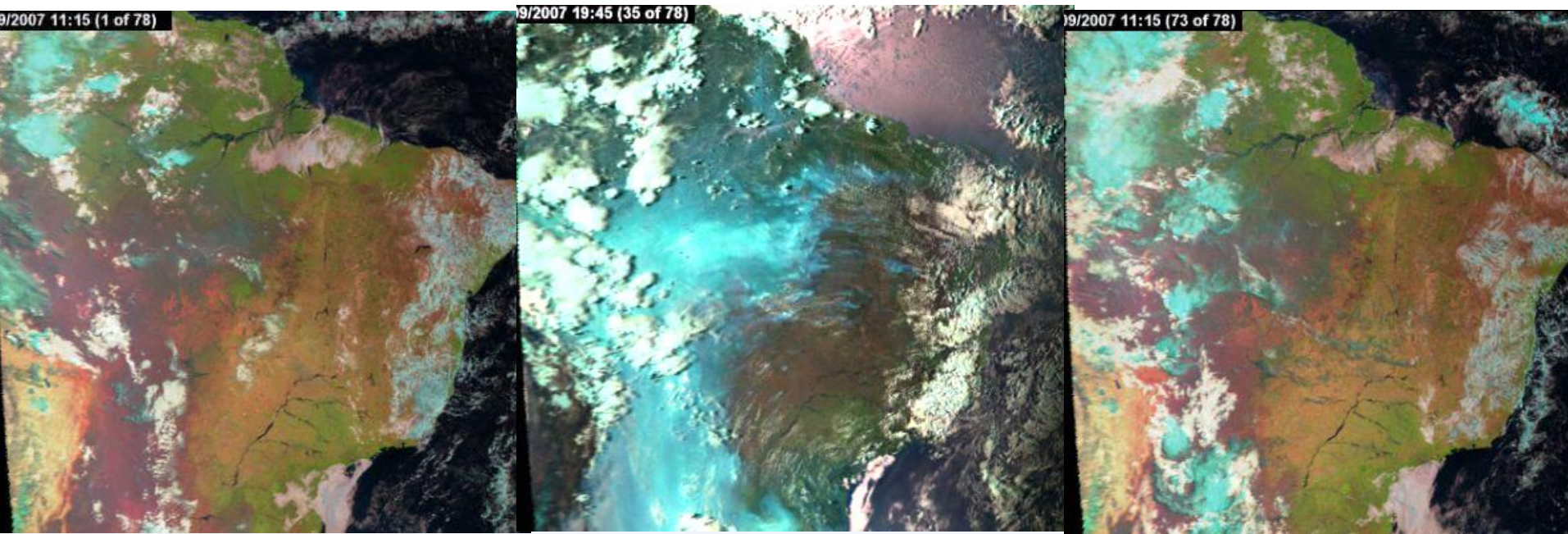
Daily Downward Surface Longwave Flux

Evapotranspiration

Evapotranspiration (30 mins)

Daily Evapotranspiration

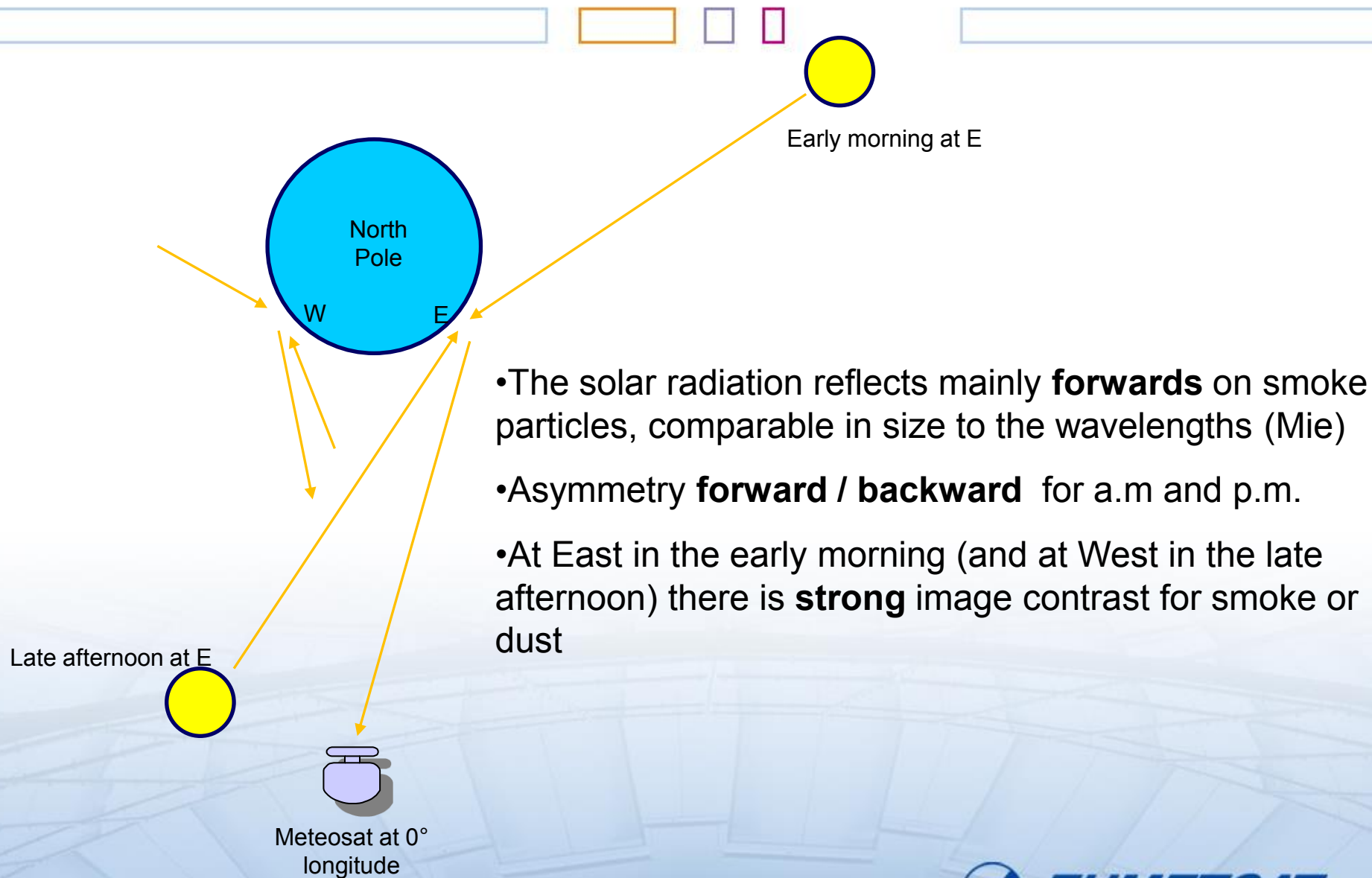
Smoke

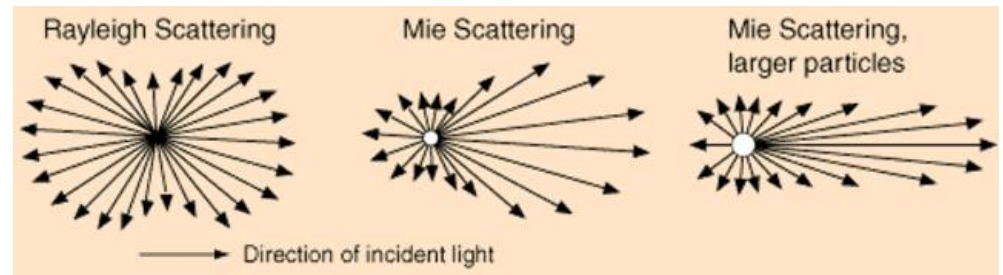
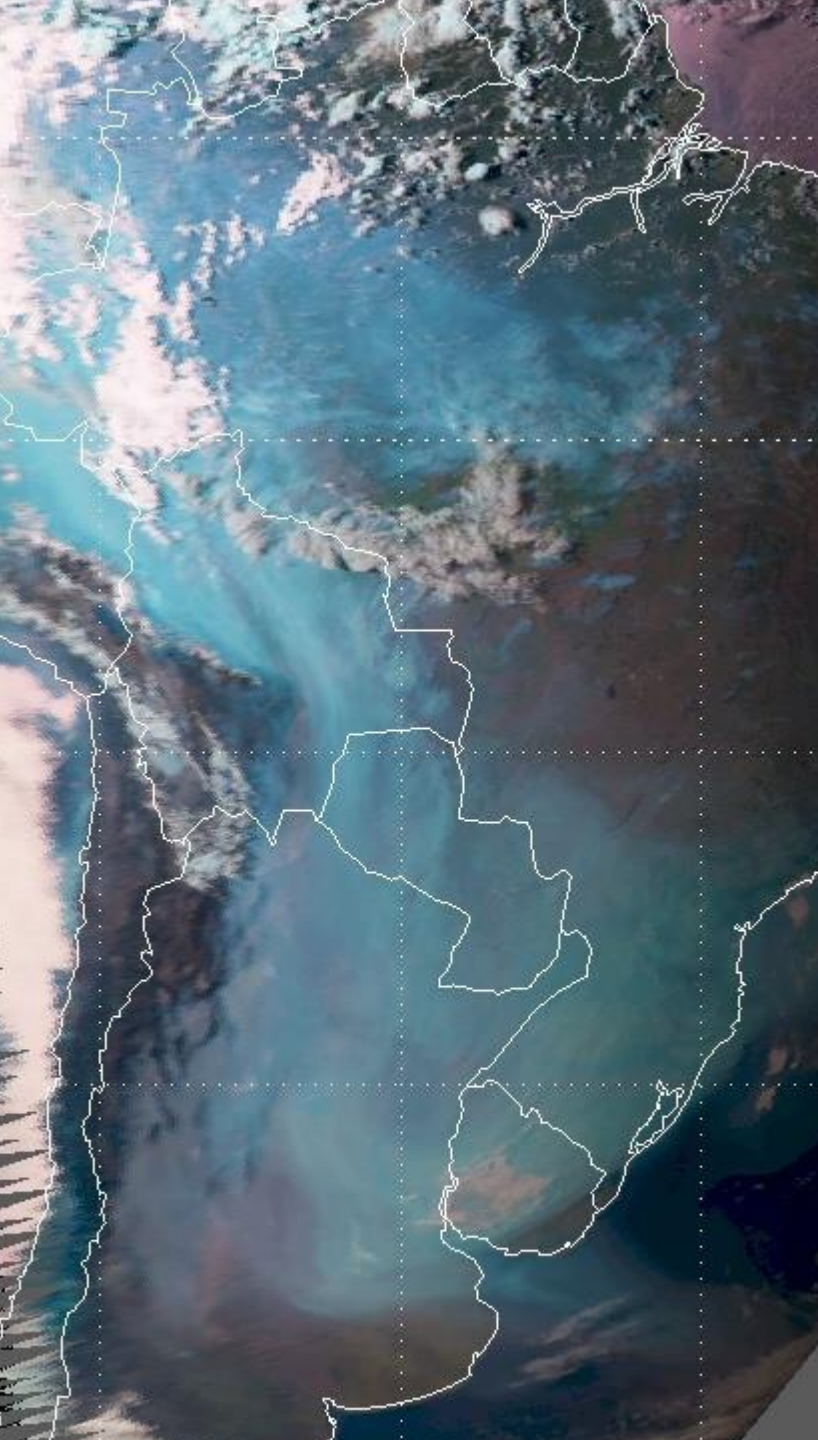


5-6 September 2007, Meteosat-9
Around sunrise and sunset times for central south America

Assuming no major smoke sink or source in 24 hours, the intensity difference is due to the sun angle

Image contrast for smoke or dust in solar images



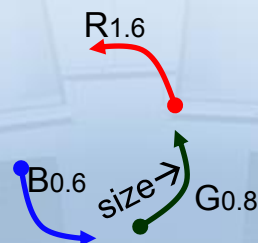
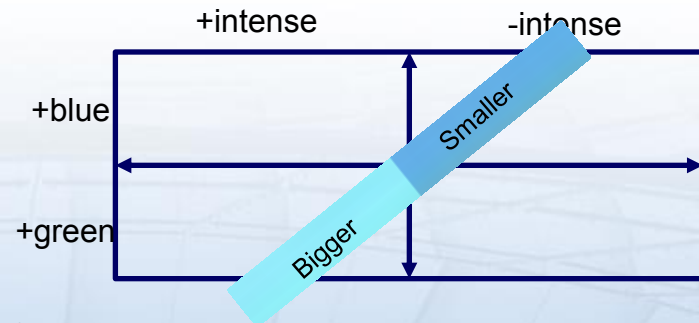


- Forward scattering favours small wavelengths (cyan)
- Bigger particles move natural RGB from blue to cyan
- Scattering intensity higher in the western late afternoon

What if smoke particles were **smaller**?

More intense scattering?

Greener or bluer in hue?



Conclusions

- Meteosat Channel 3.9 μ m is a continuous detection tool for active fires above 100m across (1 Ha), and for measuring the burnt area as reflectivities change
- **Statistics** on fires (natural or man-made) are missing or affected by sensor saturation. An approximate retrieval can be based on frequencies below saturation
- The Land SAF offers a choice of vegetation products to assess vegetation stress and **fire risk**

THANK YOU FOR YOUR
ATTENTION!

- Not in Chrome!
- Eumetsat moodle, might request user name (free)



Brightness temperature at 3.9 micron for detecting fire in the pixel

Fires brightness temperature (BT)

This applet interface describes with sliders the characteristics of the atmosphere, and provides the BT at 3.9 μ m (or 10.8 μ m with the button) for different types of soil, times of the day (sun elevation), intense and extensive fires and cloud above.

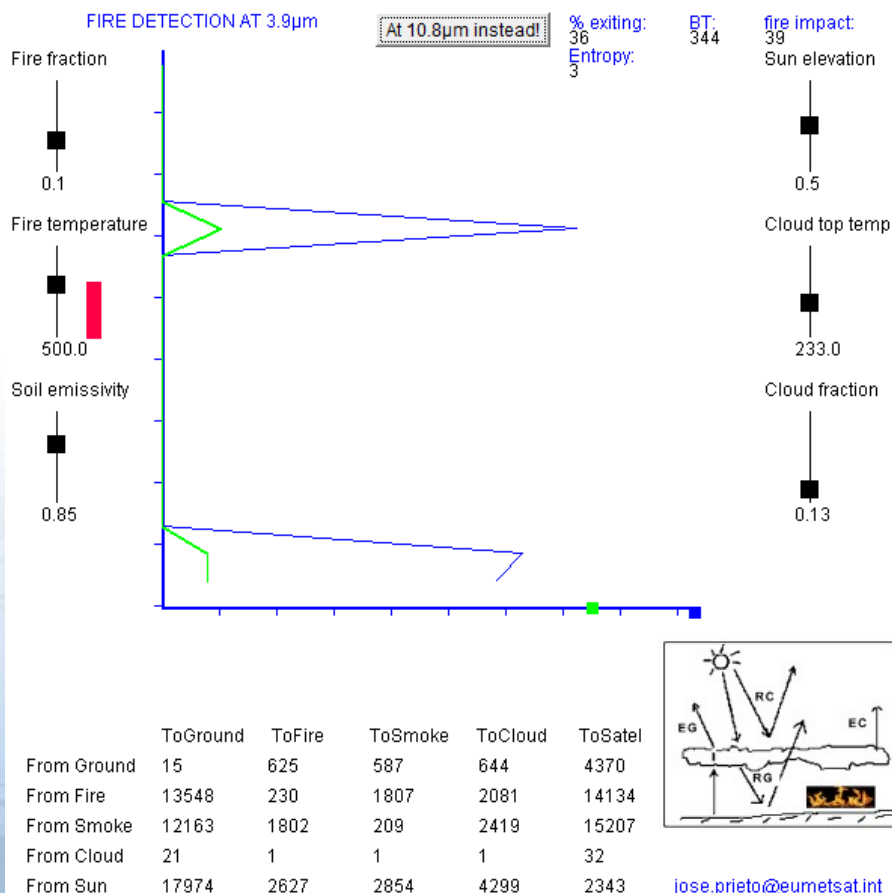
Green indicates the amount of cloud, fire and soil emissivity (bottom green square). Blue is the fraction of radiation from different sources exiting the Earth, the rest being absorbed by the Earth-atmosphere.

The global value is given in upper line. The reddish square close to the Fire Temperature slider gives an idea of the extent and intensity of the fire and its size depends on the first two sliders. Its colour is given by the brightness temperature difference (fire impact) with and without the fire. Red indicates that the difference is above 5 K, so the satellite could see the fire.

The bottom table specifies the actual amounts of energy exchanged by the elements in the scene, relative to 100000 photons emitted in total by all surfaces, upwards and downwards.

Back to work:

1. First, set all sliders to 0, but "Soil emissivity" to 1 (absolutely non-reflective ground. Usually, it should be between 0.50 for desert or savannah and 0.85 for thick forest). Notice 100% of the emitted radiation reaches the satellite. This proportion will decrease when new sources are added.
2. Set the Cloud thickness to some intermediate value, and observe the changes. What do you expect for a brightness temperature, as a function of the Cloud top temperature? Are you correct? If not, why?
3. Back to Cloud thickness zero, try with Sun elevation, the sun rising over the horizon and sending radiation at 3.9 μ m into the atmosphere and back to the satellite. Any changes when you move the slider? How does BT vary when we change Soil emissivity on the ground?



Did you pay attention?

- Fire analysis from satellites is a complex matter
 - Yes
 - No
- You might confuse a fire pixel at $3.9\mu\text{m}$ with ...
 - Thin cirrus during the day
 - A water surface affected by sun glint
 - Ash particles in the atmosphere
- Smoke is better detected by solar channels when the satellite is...
 - In the backward scattering direction of sun radiation
 - In the foreward scattering direction of sun radiation
- The main obstacle to assess fire characteristics with Meteosat $3.9\mu\text{m}$ is..
 - Channel saturation at a low temperature
 - Smoke and cloud generated by the fire
 - Uncertainty on fire size
 - CO_2 generated by the fire